

# TIDAK: the updated Suhora database of O–C diagrams and elements of eclipsing binary stars

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The database of minima, O–C diagrams, and linear elements of eclipsing binary stars maintained at the Mt. Suhora Observatory of the Pedagogical University of Krakow in Poland is widely referenced in eclipsing binary studies. Here we present the latest version of the database, along with a new, graphical interface including features such as interactive O–C diagrams. This is the final result of an extensive modernization program aimed at making the database more useful and convenient for researchers and observers.

## 1 Introduction

Observations of the times of minima of eclipsing binary stars have long been an important tool for the study of these types of objects. Comparison of the observed times of minima with those predicted from an ephemeris (often presented as ‘O–C’, i.e. ‘observed – calculated’, values plotted on an ‘O–C diagram’) can reveal physical effects including stellar evolution, mass transfer between the components, or gravitational interaction with a third star or, potentially, planets. For this kind of analysis, it is beneficial to have records of minima times that cover as long a time as possible, in order to resolve slow changes in the binary system.

## 2 Previous editions of the database

For several decades through the work of Prof. J. M. Kreiner, times of minima have been collected at the Mt. Suhora Observatory of the Astronomy Department of the Pedagogical University in Krakow base and used to generate up-to-date ephemerides and O–C diagrams of eclipsing binary stars. Previously the ephemerides and O–C diagrams were presented in the *Atlas of O–C Diagrams* (Kreiner et al., 2001), available as a set of physical books and also online on the website of the Astronomy Department of the Pedagogical University, and subsequently only as an online service (Kreiner, 2004). The collected times of minima were also available to be sent to researchers on request.

However, in recent years, and especially with the growth of automated surveys, the number of stars and minima times had outgrown the technical capabilities of the underlying database. It became necessary to replace the database architecture, and at the same time, the opportunity was taken to create a new online interface with more features.

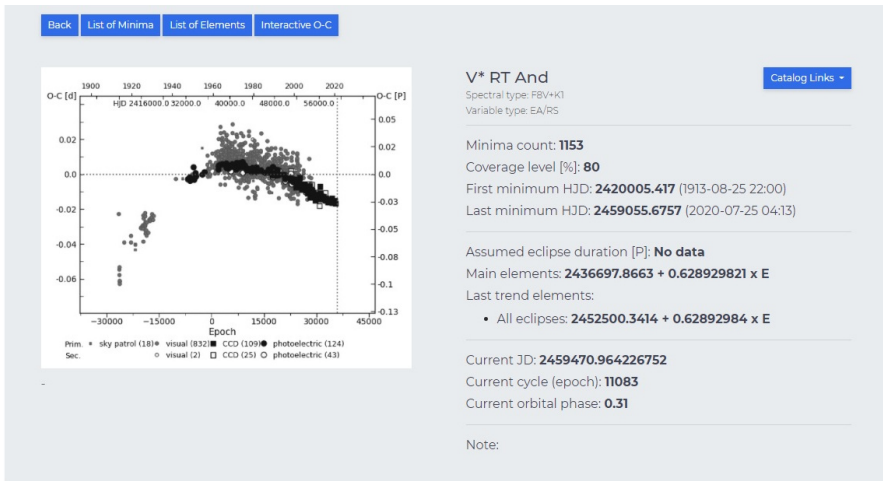
Select  
AND - Andromeda

Show  entries

Search:

Star	CCVS No	RA	DEC	P [d]	All Ecl	Pri Ecl	Sec Ecl	Var Type	Last Trend	Mag	Last Obs Year	Tags
V* RT And	10012	23:11:09.981	+33:01:33.031	0.62893	1153	1083	70	EA/RS	A	B 9.55	2020	-
V* SV And	10025	00:18:16.3767	+43:42:41.44	54.30890	37	36	1	EA	A	B 10.85	2016	-
V* TT And	10027	23:12:29.034	+46:08:51.079	2.76331	105	105	0	EA	A	B 11.75	2018	-
V* TW And	10030	00:03:18.2272	+32:50:45.102	4.12276	195	193	2	EA	A	B 9.46	2018	-
V* LU And	10034	00:43:45.0803	+30:56:19.064	1.48630	287	271	16	EA	A	B 12.46	2020	-
V* WW And	10045	23:44:53.5469	+45:41:11.640	23.28524	31	29	2	EA	A	B 11.41	2005	Observed
V* WX And	10046	00:42:37.3404	+38:44:53.981	3.00718	43	43	0	EA	A	B 13.05	2016	-
V* WZ And	10048	01:01:43.6466	+38:05:46.489	0.69566	408	378	30	EB	A	B 11.84	2020	-

Fig. 1: List of eclipsing binaries in the constellation Andromeda.



Computed times of minima

Show  entries

Search:

Date [UT]	HJD	Time to [h]	Type
2021-09-13 06:29:58	2459470.770817	-4.66	Pri
2021-09-13 14:02:48	2459471.085282	2.89	Sec
2021-09-13 21:35:38	2459471.399747	10.43	Pri

Fig. 2: Star overview page for RT Andromedae.

### 3 The new version: TIDAK

The new version of the database (stylized as ‘TIDAK’, for ‘Timing Database at Kraków’) uses PostgreSQL to store the data with an interactive website created using Django and Python. The user can find stars by searching for catalog identifiers or coordinates, with the SIMBAD service (Wenger et al., 2000) used to interpret the query, or by constellation (e.g. Fig. 1). The individual overview page (Fig. 2) shows an O–C diagram, the ephemeris elements, and a list of upcoming minima to

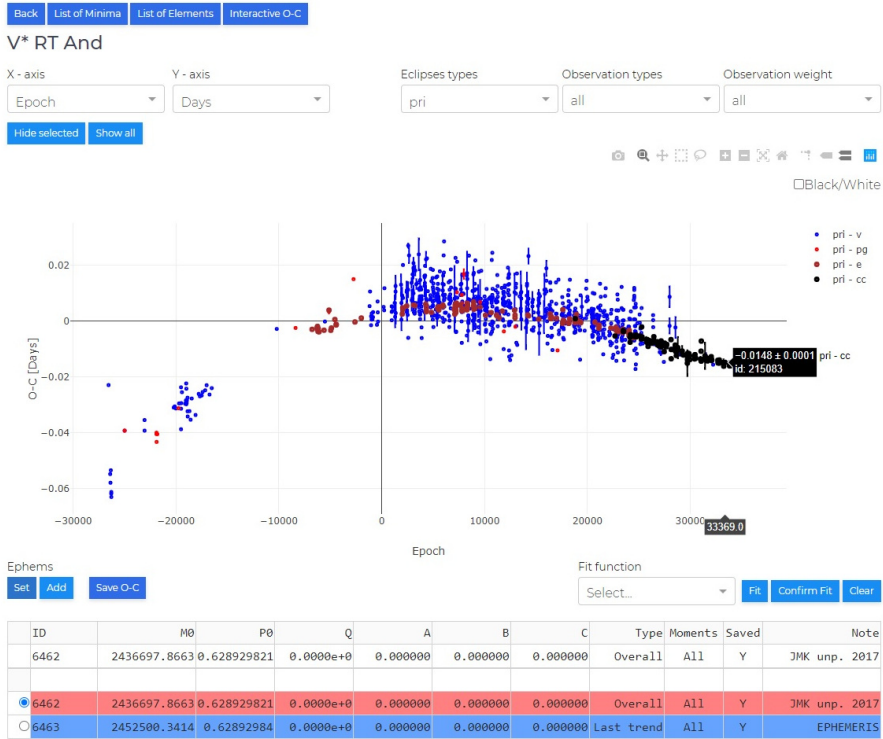


Fig. 3: Interactive O–C plot.

aid with planning observations.

New features include an interactive O–C plot (Fig. 3) where the user can explore different ephemerides and fit new elements, unlimited ephemerides for each star (previously only 3 ephemerides were stored) allowing more flexibility for analyzing trends and changes in behavior, and the possibility for users to download lists of minima and O–C values directly for themselves.

## 4 Summary

The database currently contains data for 8054 stars with 254 thousand minima, gathered from literature and direct observations by the staff of the Mt. Suhora Observatory. New minima are continuously being added. The database can be accessed at <https://www.as.up.krakow.pl/tidak>.

## References

- Kreiner, J. M., *Acta Astron.* **54**, 207 (2004)
- Kreiner, J. M., Kim, C.-H., Nha, I.-S., *An Atlas of O-C Diagrams of Eclipsing Binary Stars* (2001)
- Wenger, M., et al., *A&AS* **143**, 9 (2000)