

# Main Sequence to Starburst Transitioning galaxies

GRB hosts at redshift  $\sim 2$

(ApJ, 2023, 952, 125N)



**Jakub Nadolny**



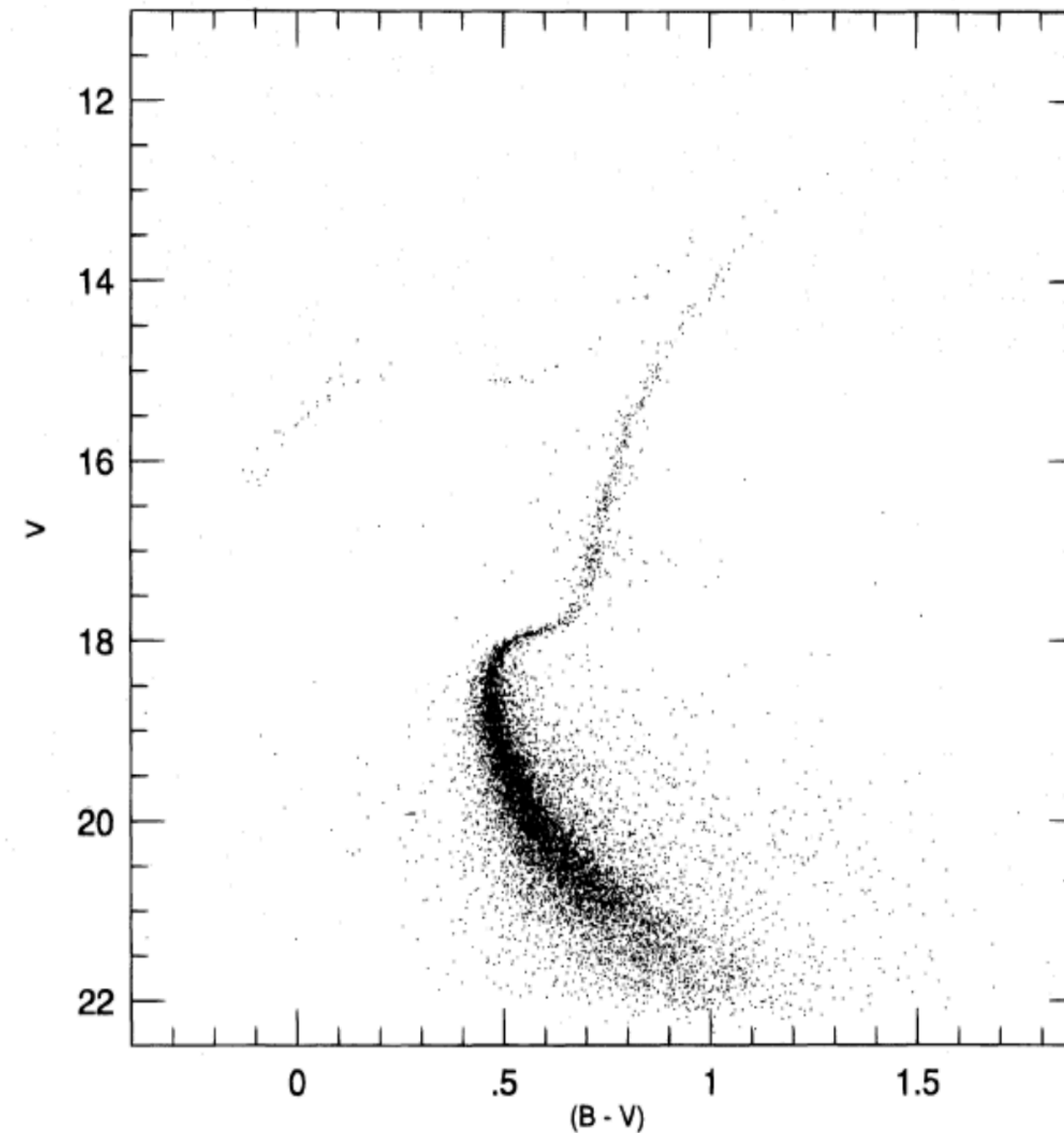
Michał Michałowski, Ricard Rizzo, Agata Karska, Jesper Rasmussen,  
Jesper Sollerman, Jens Hjort, Andrea Rossi, Martin Solar, Radosław  
Wróblewski, and Aleksandra Leśniewska

# Layout

- 1) Main Sequence**
- 2) Gamma-ray burst**
- 3) Motivations**
- 4) Results**

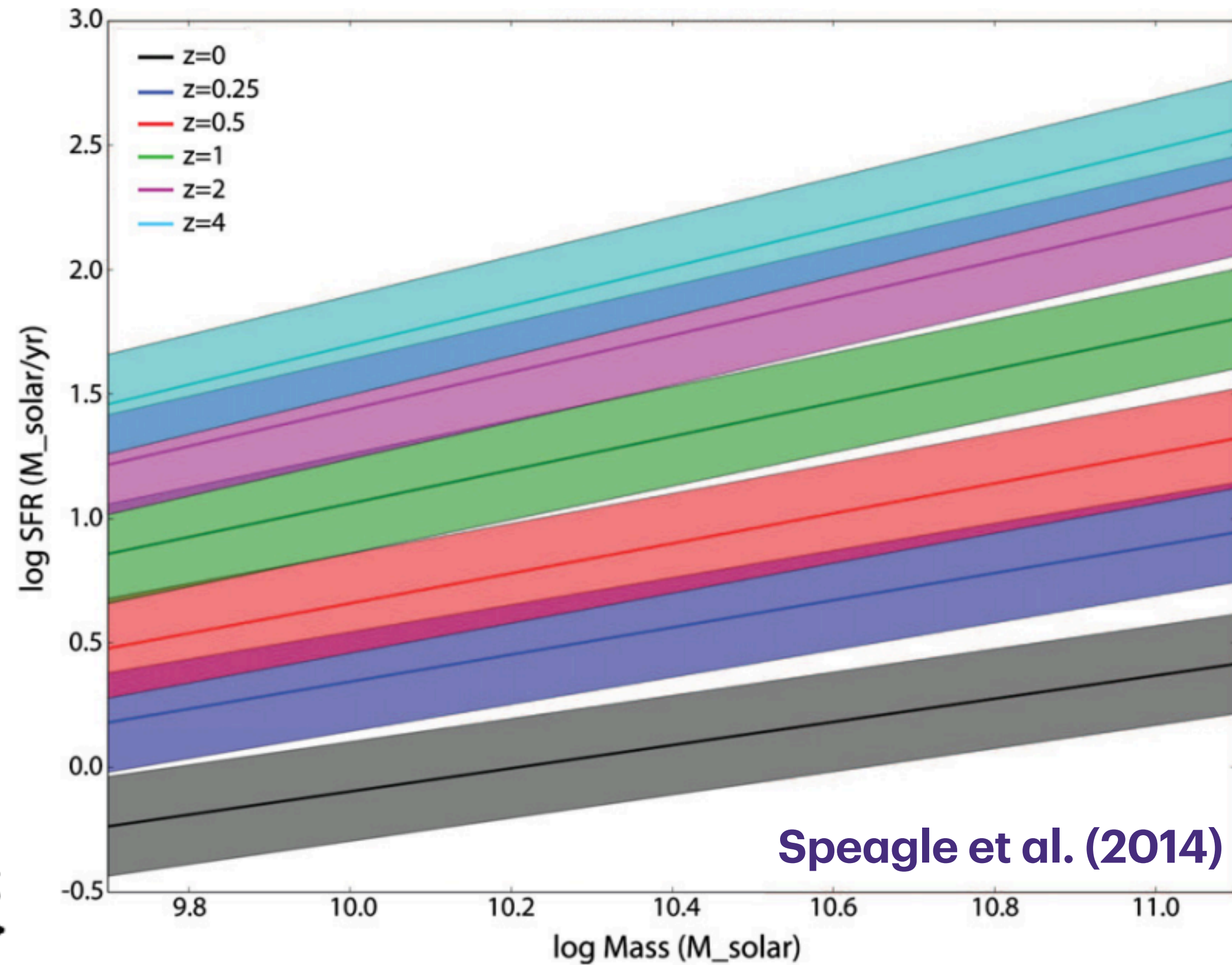
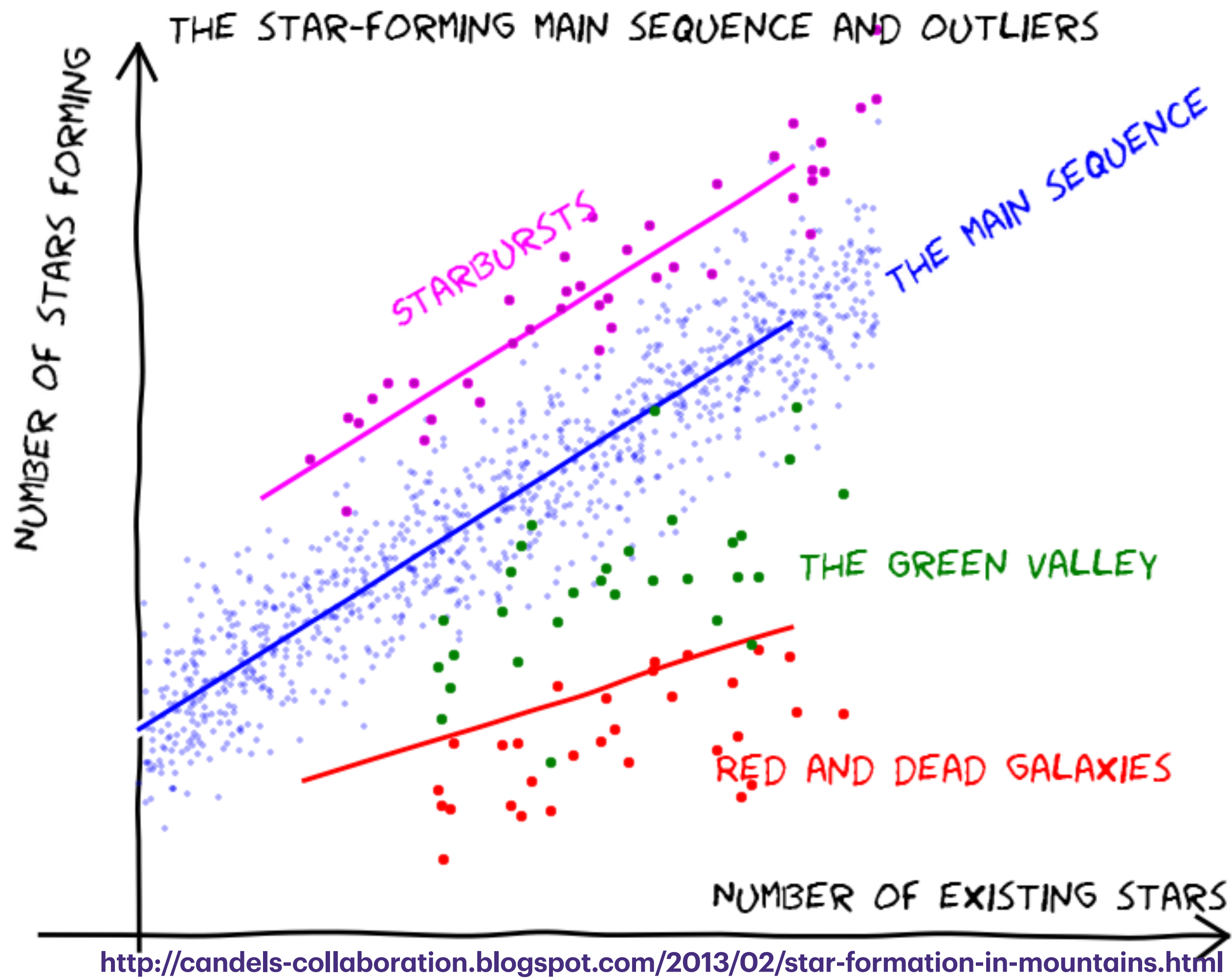
# Main Sequence

of Star Forming Galaxies (not Main Sequence Stars)

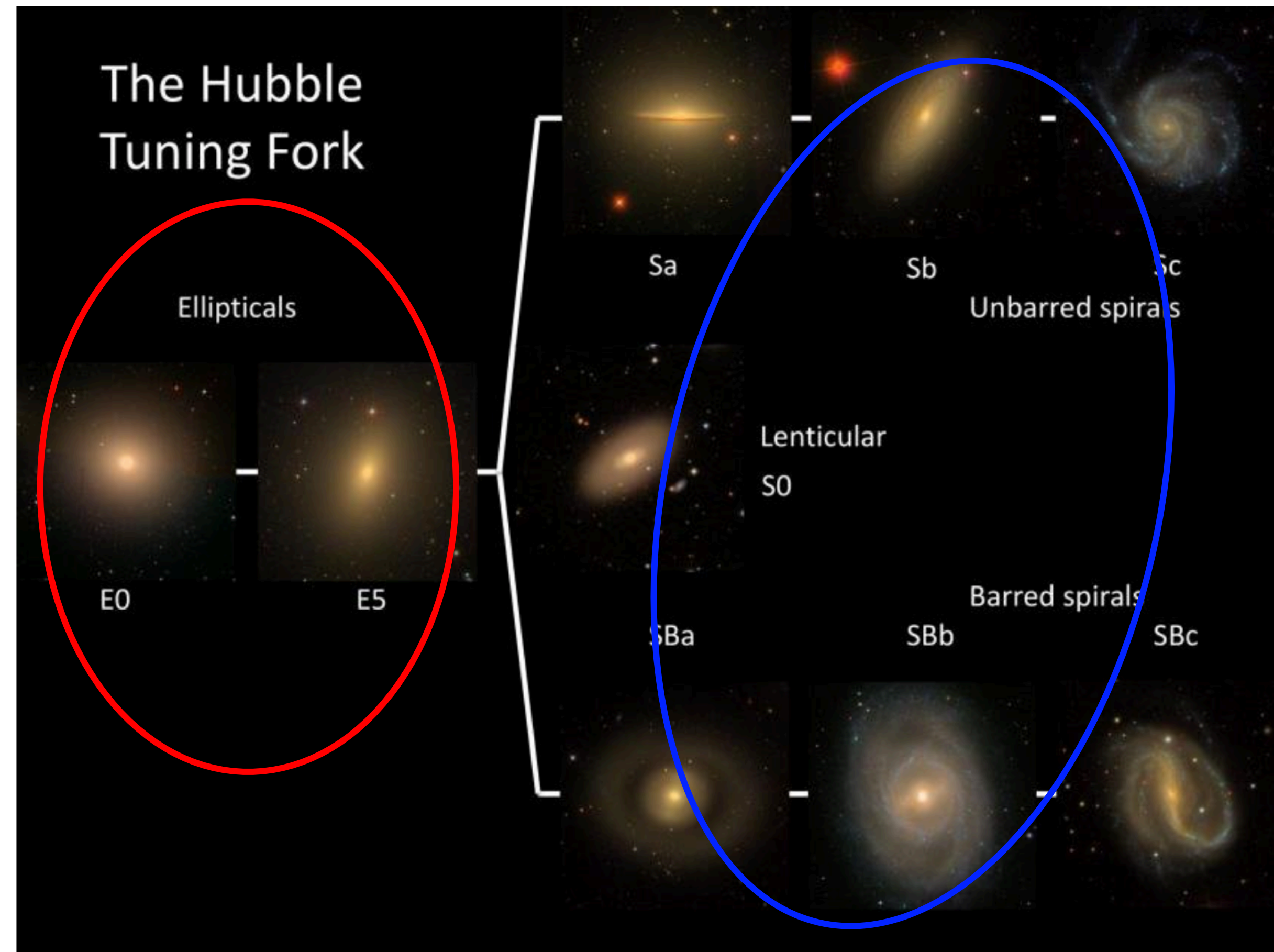
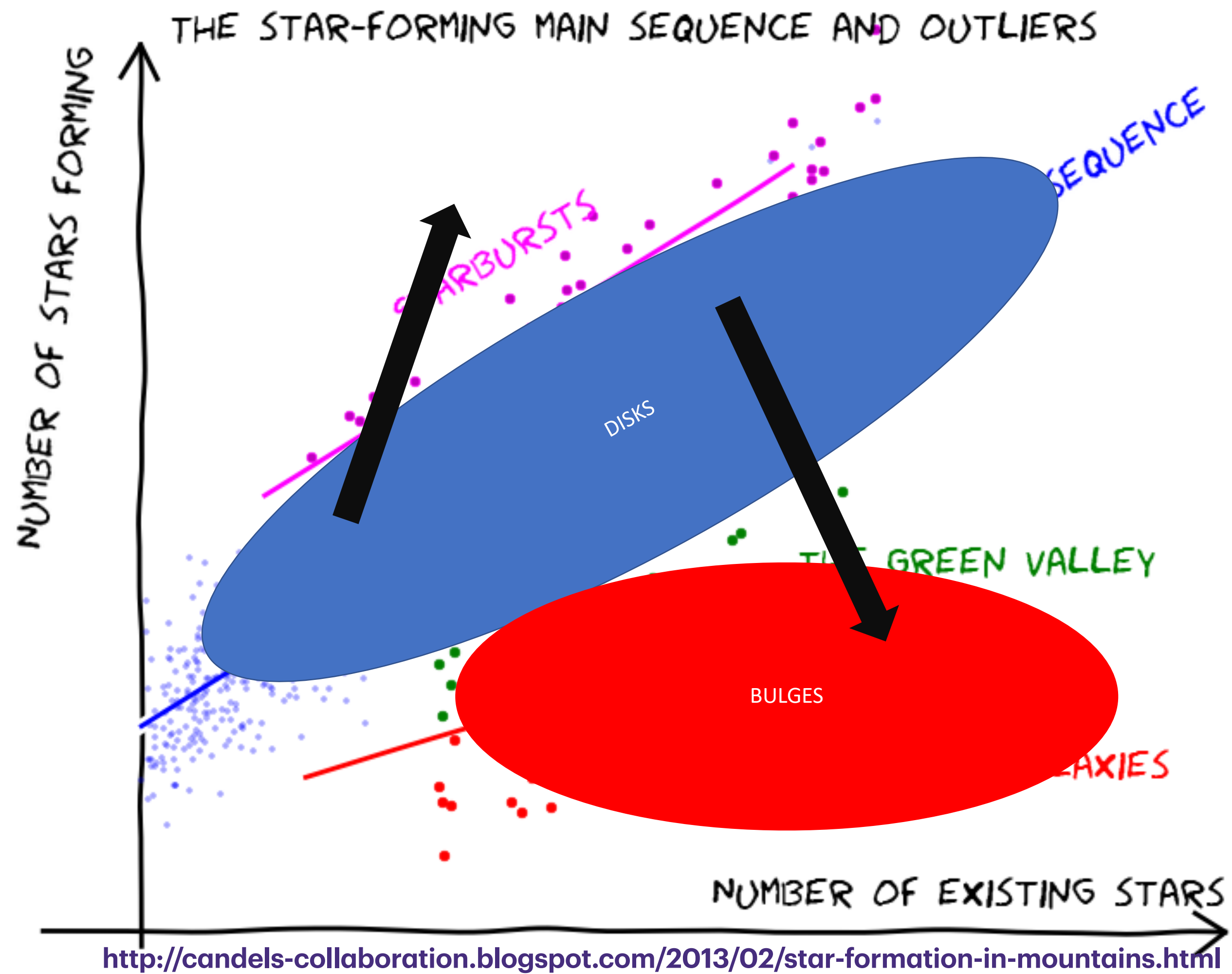


CMD (Sandquist et al. 1996) of M5 (NGC 5904)

# Main Sequence



# Main Sequence

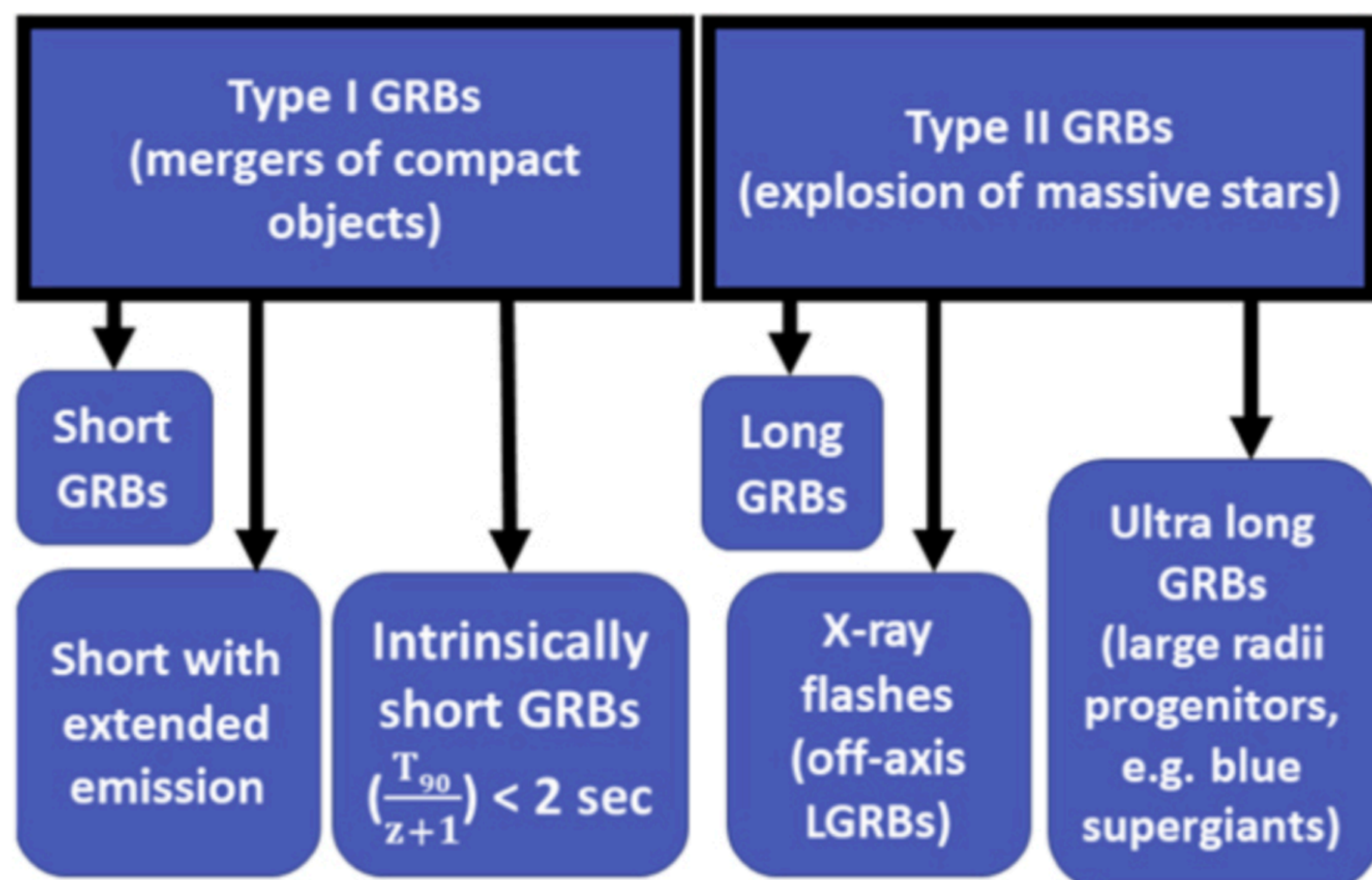


# Gamma-ray burst

Gerardo and Aditya yesterday talks

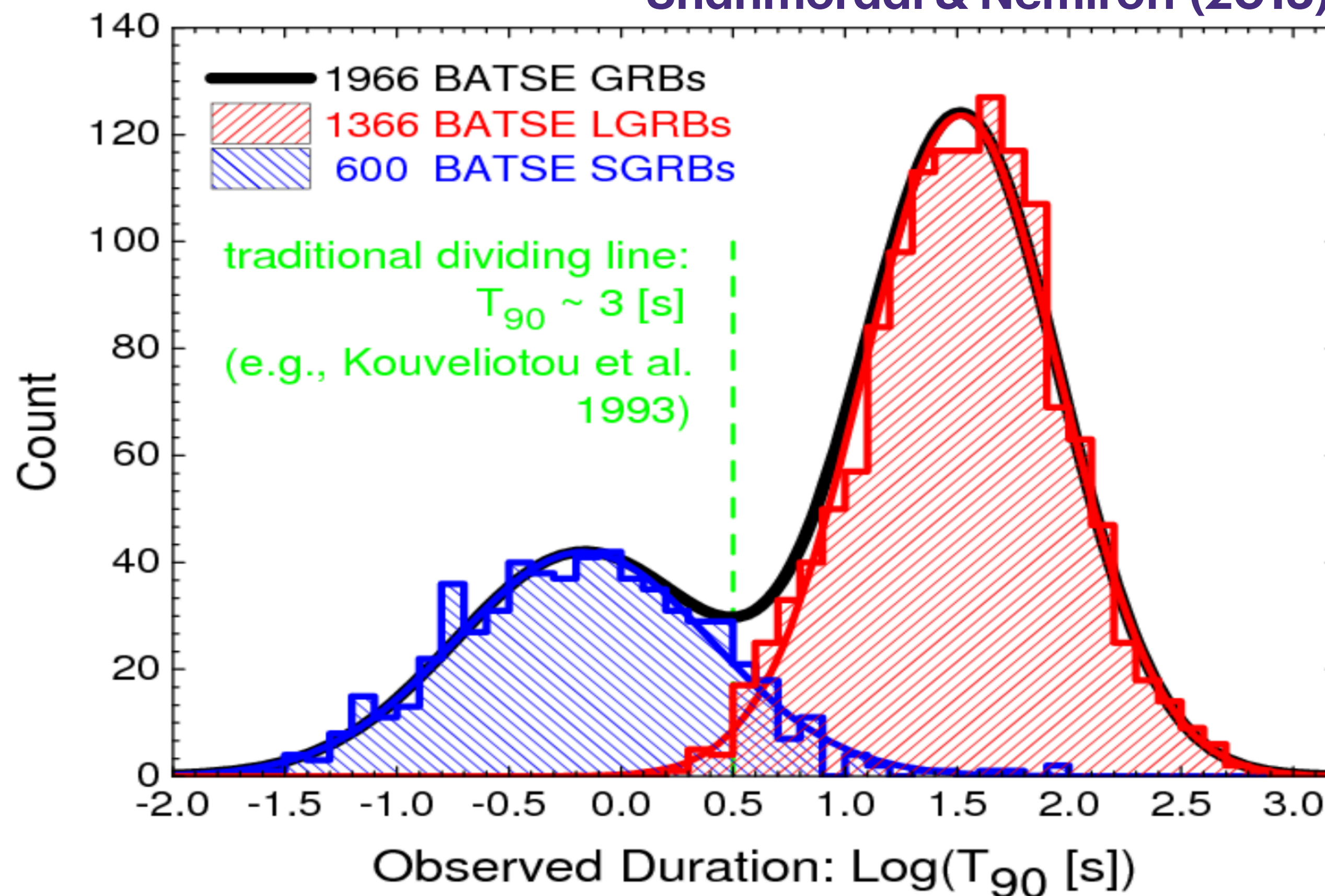
**Short GRB: NS-NS/NS-BH**

**Long GRB: SNe explosion**



Maria G. Dainotti et al (2023)

Shahmoradi & Nemiroff (2015)

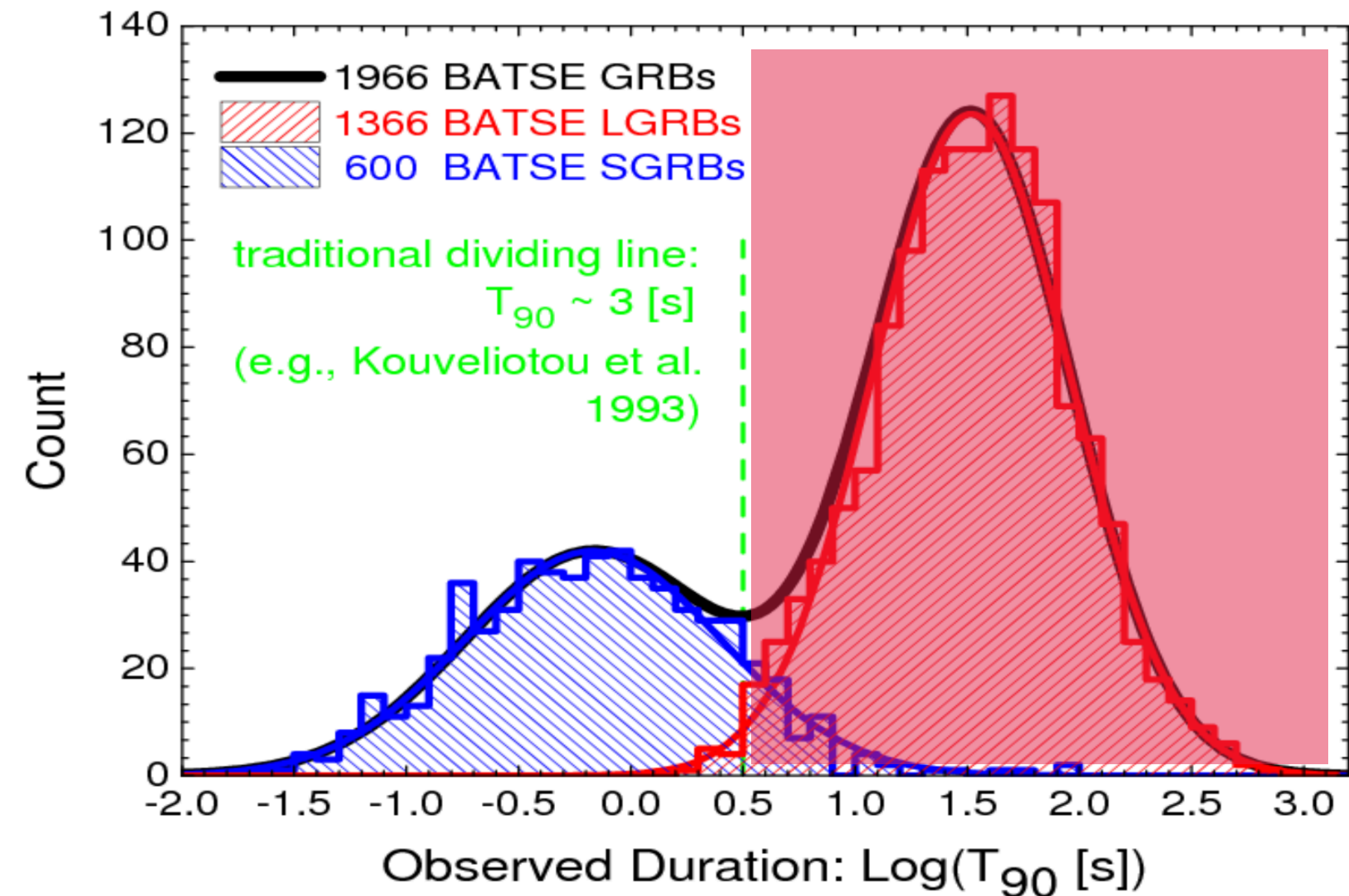
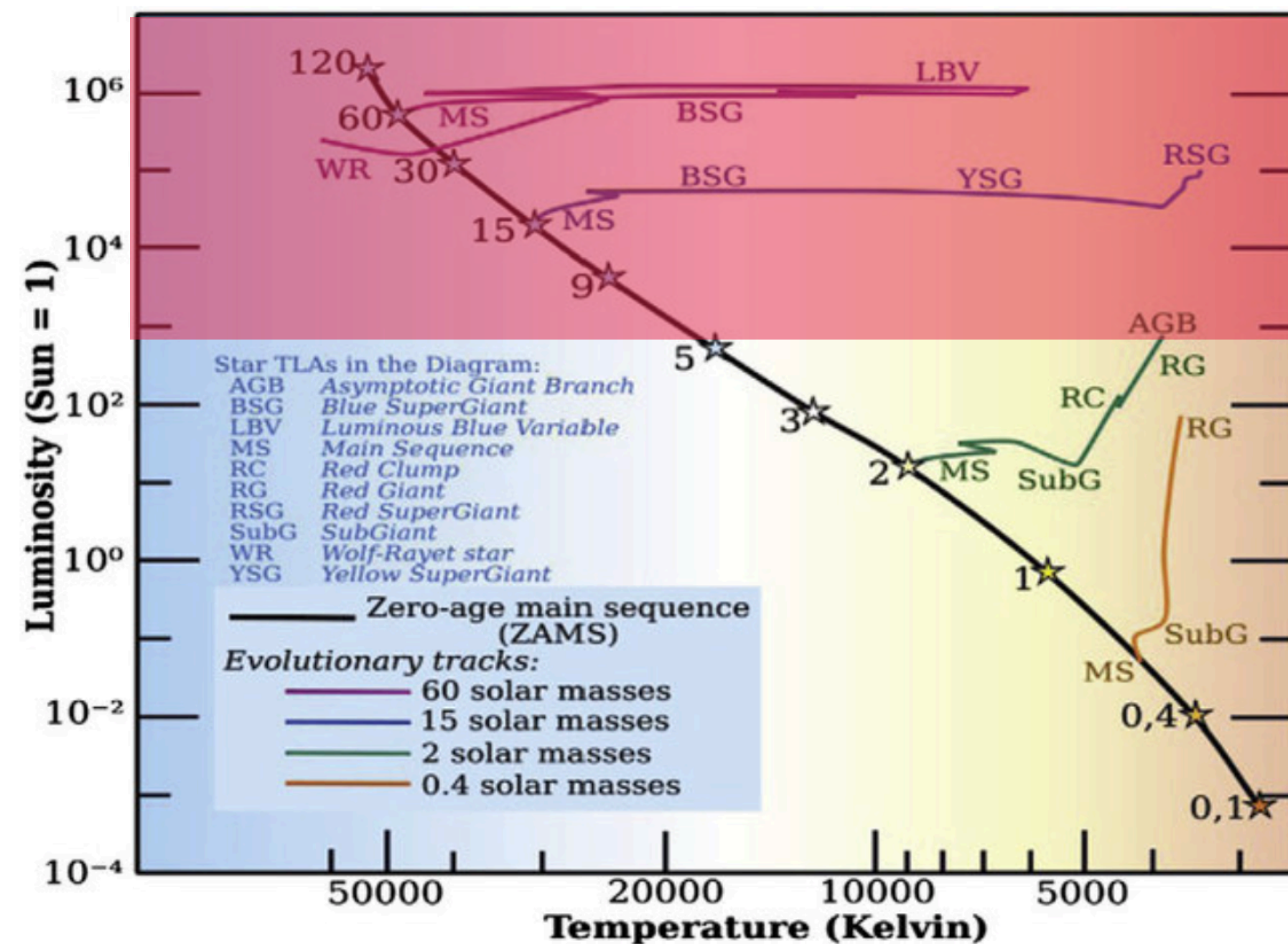


# Gamma-ray burst

## Chasing recent star-formation in GRB host galaxies

Long GRBs originate in explosions of short-lived massive (>8 Msun) stars

Hjort+03, Stanek+03

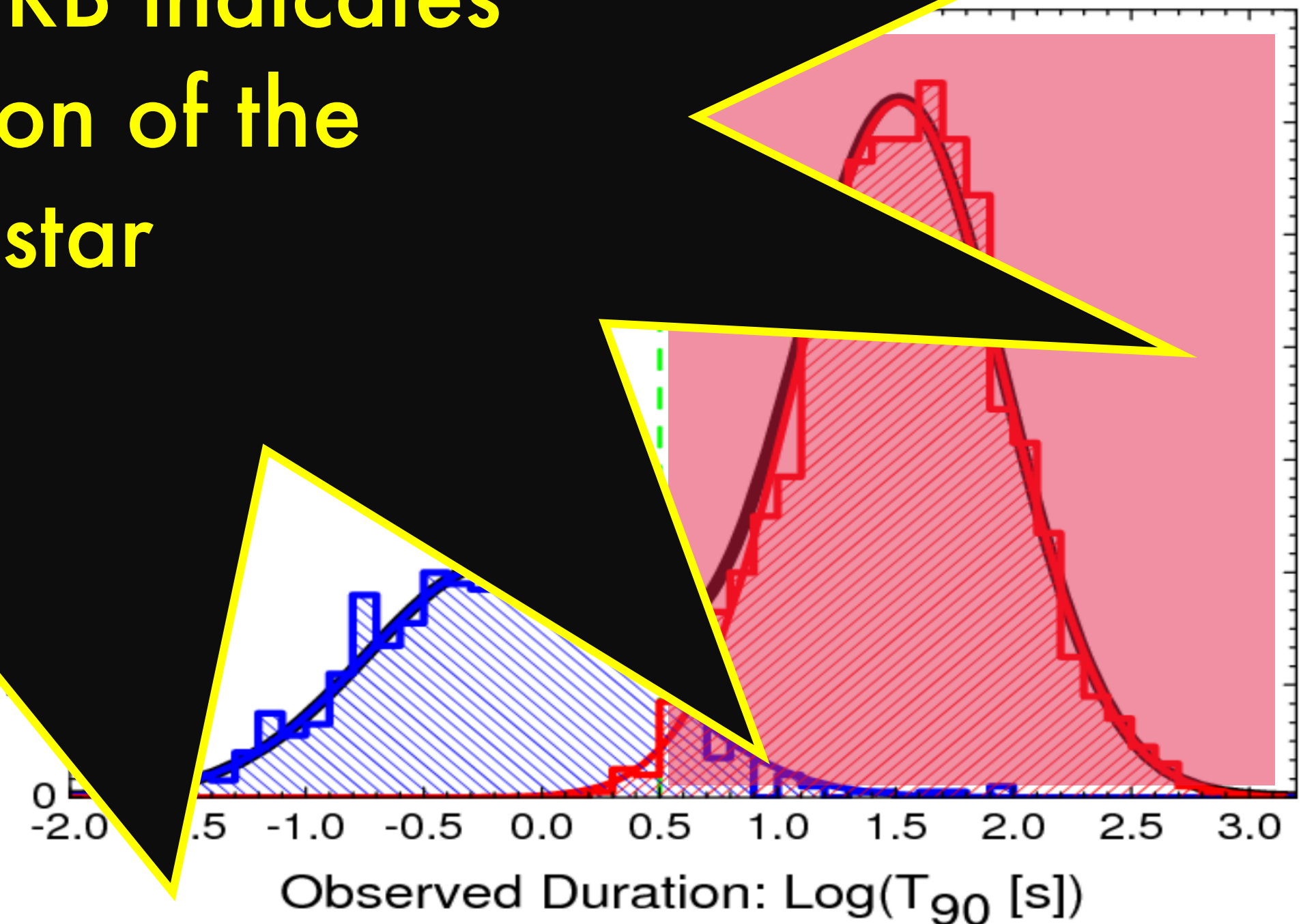
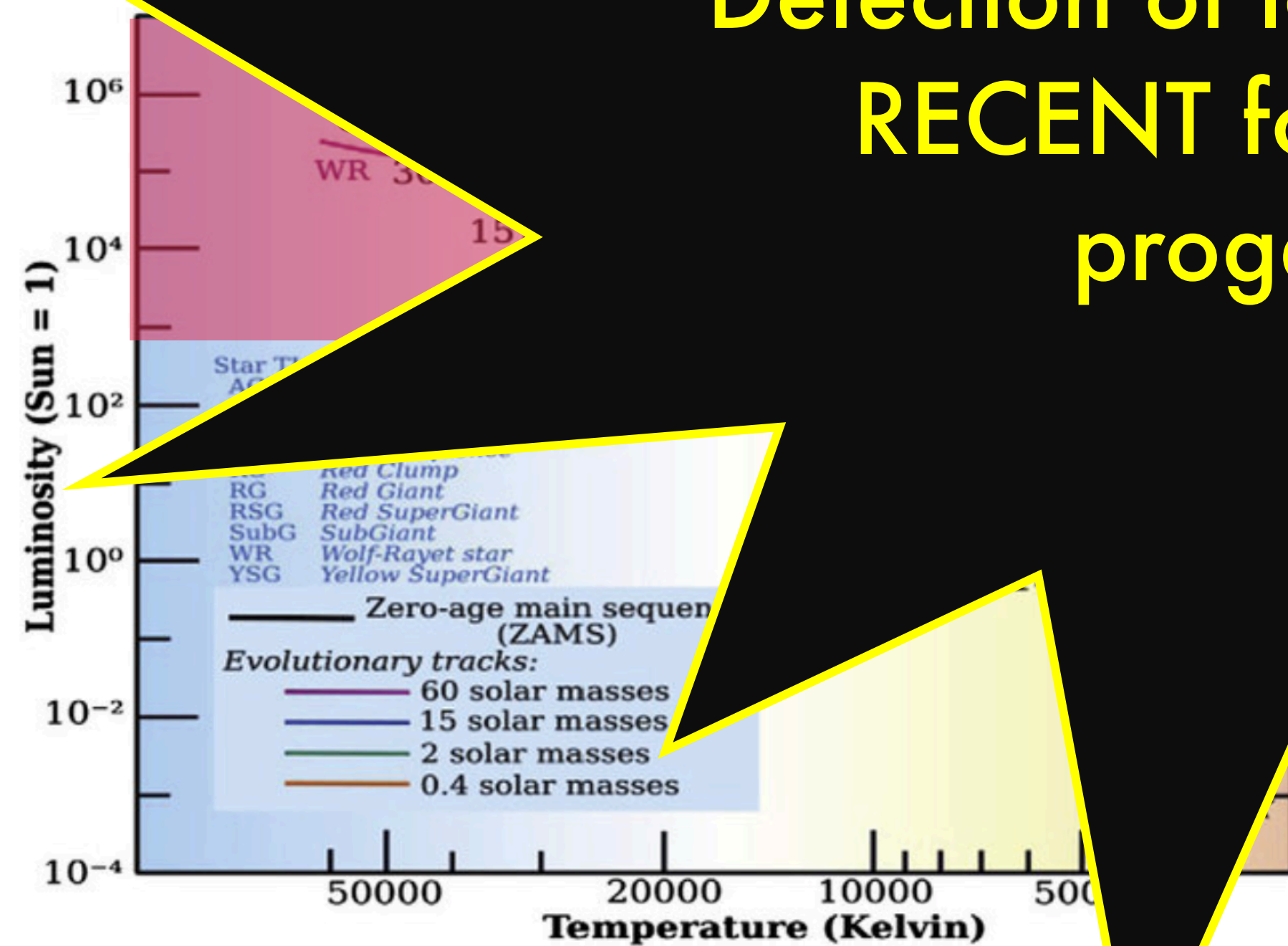


# Gamma-ray burst

Characterizing recent star formation in GRB host galaxies

Long GRBs originate from the death of massive (>8 Msun) stars

Detection of long GRB indicates  
RECENT formation of the  
progenitor star





# Motivations

- 1) To test whether GRB hosts are similar to other galaxies**
- 2) How the GRB, gas content and SFR are connected**

# Sample Selection

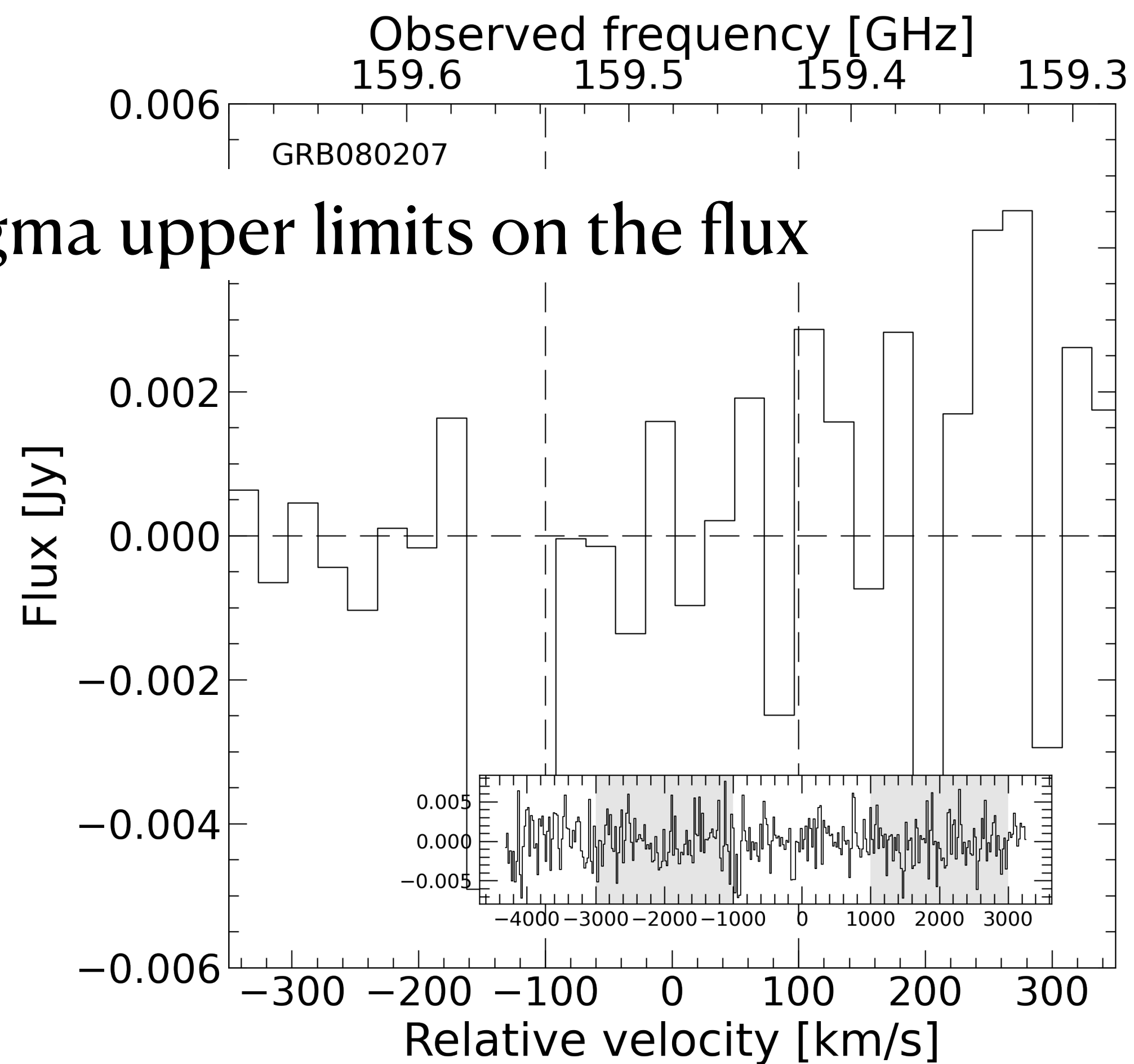
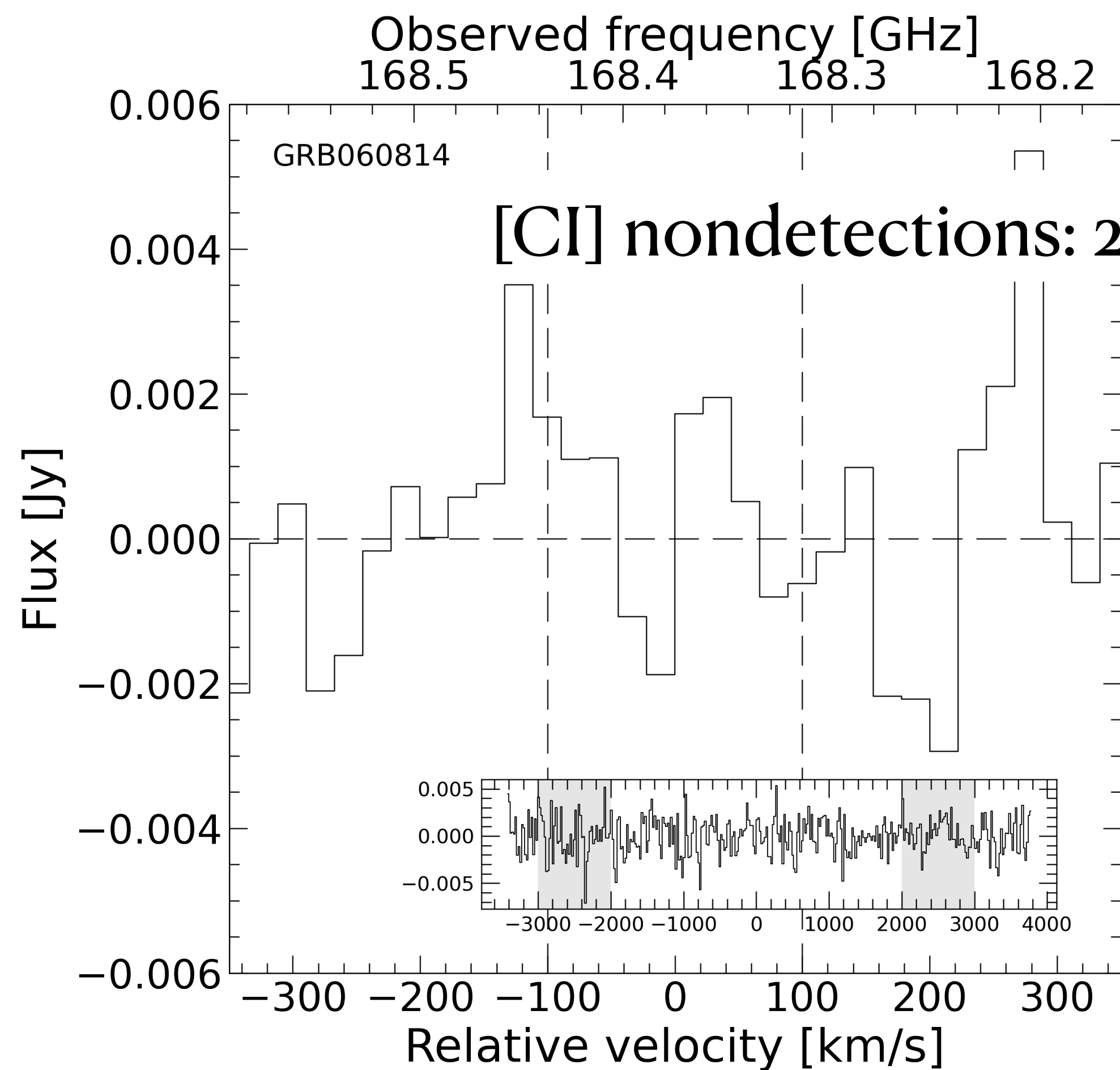
- 1) Availability of IR or radio detections: SFR estimators
- 2) Spectroscopic Redshift: to assure that neutral carbon [C I](1-0) could be observed

- 1) Seven potential targets  
(GRB ~~051006~~, ~~051022~~, **060814**, **061121**, **080207**, **100316D**, and ~~111005A~~)
- 2) Two telescopes: **APEX**, and **IRAM**.
- 3) Between 13h and 17h on-source.

**APEX: GRB 061121 and 100316D — upper limits not sufficiently constraining.**

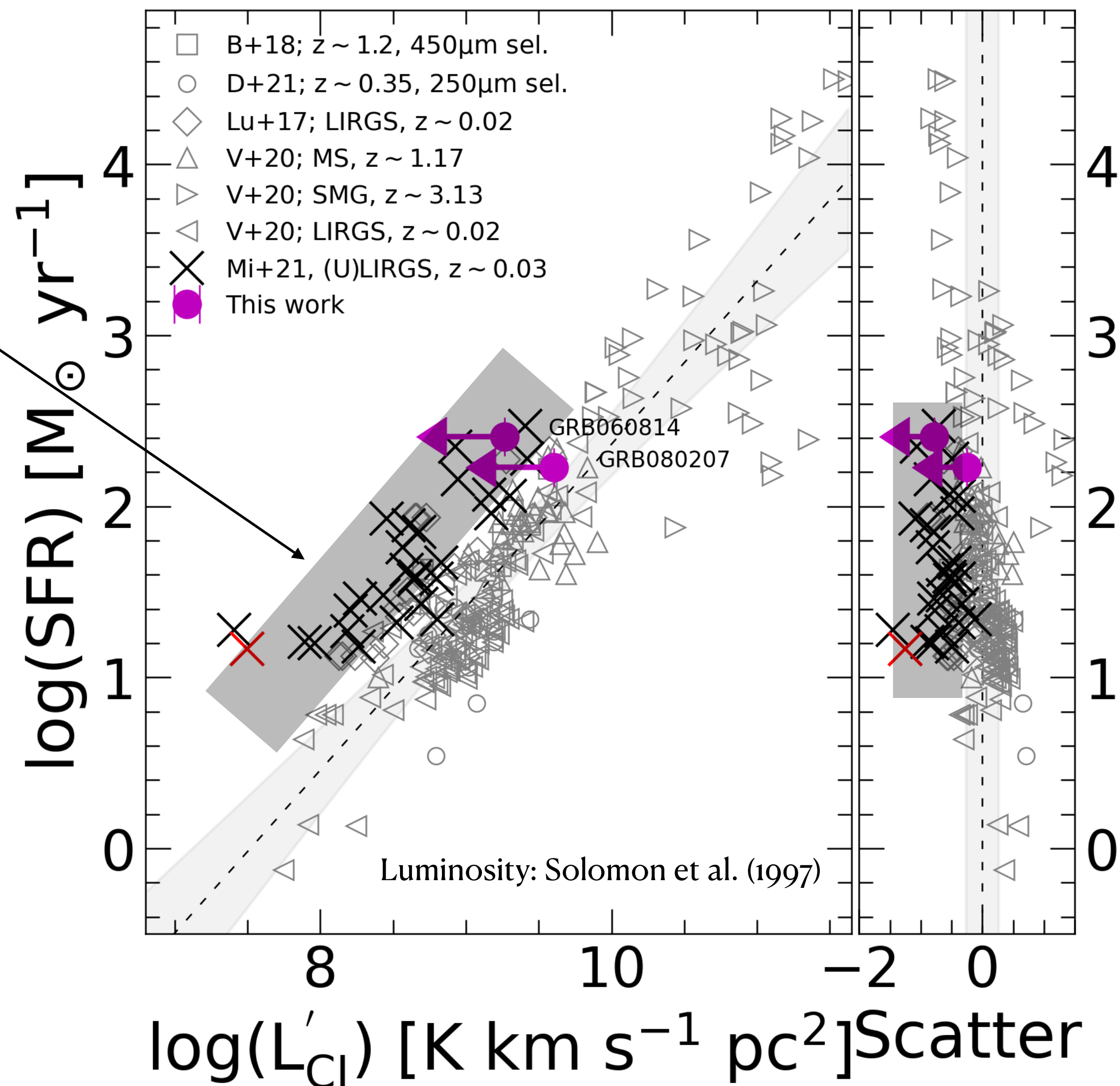
# Results

## IRAM: GRB 060814 and 080207



# Results

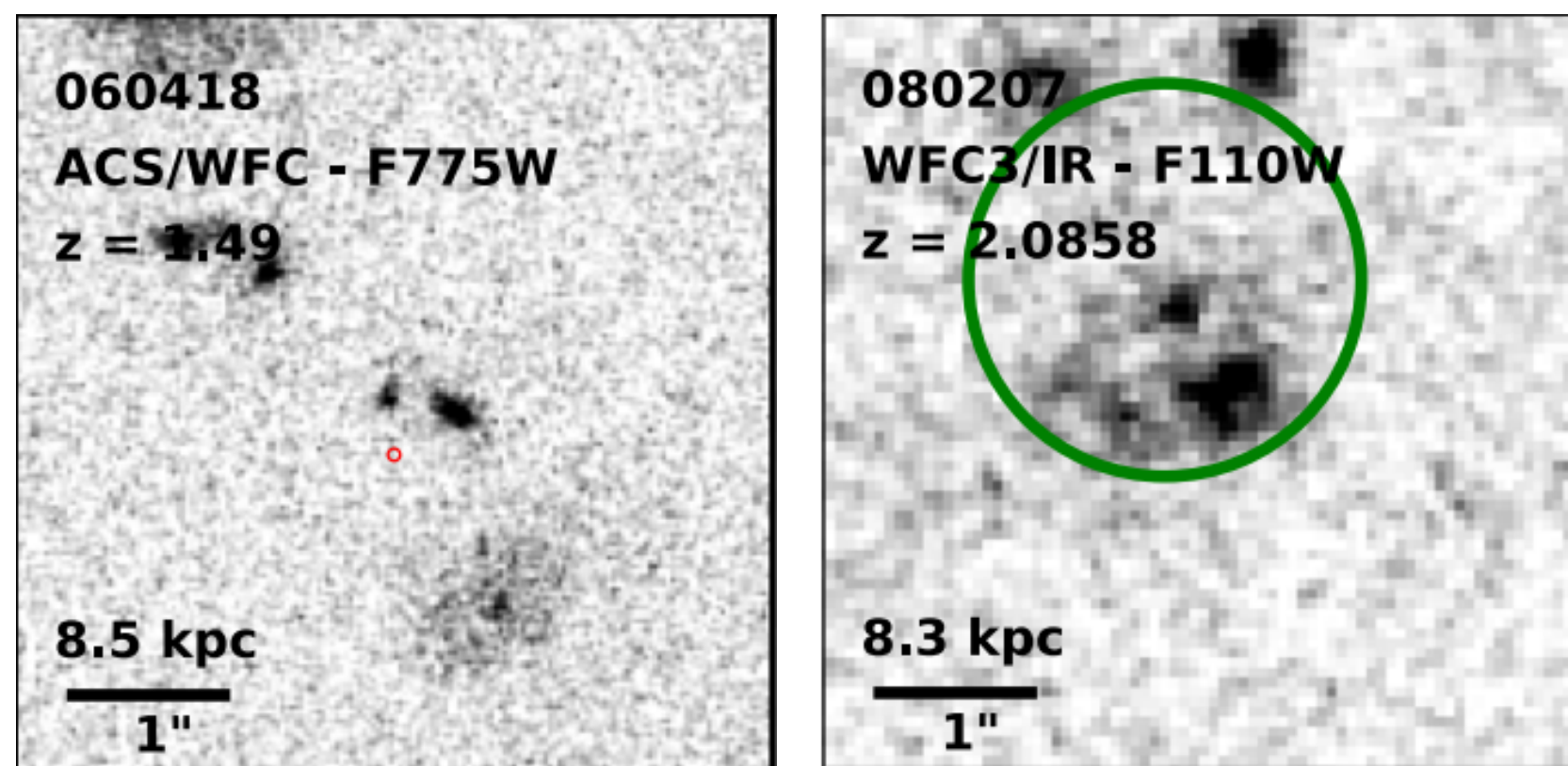
[CI] nondetections:  
upper limits  $L[CI]$  — similar to local  
*merging* ULIRGS



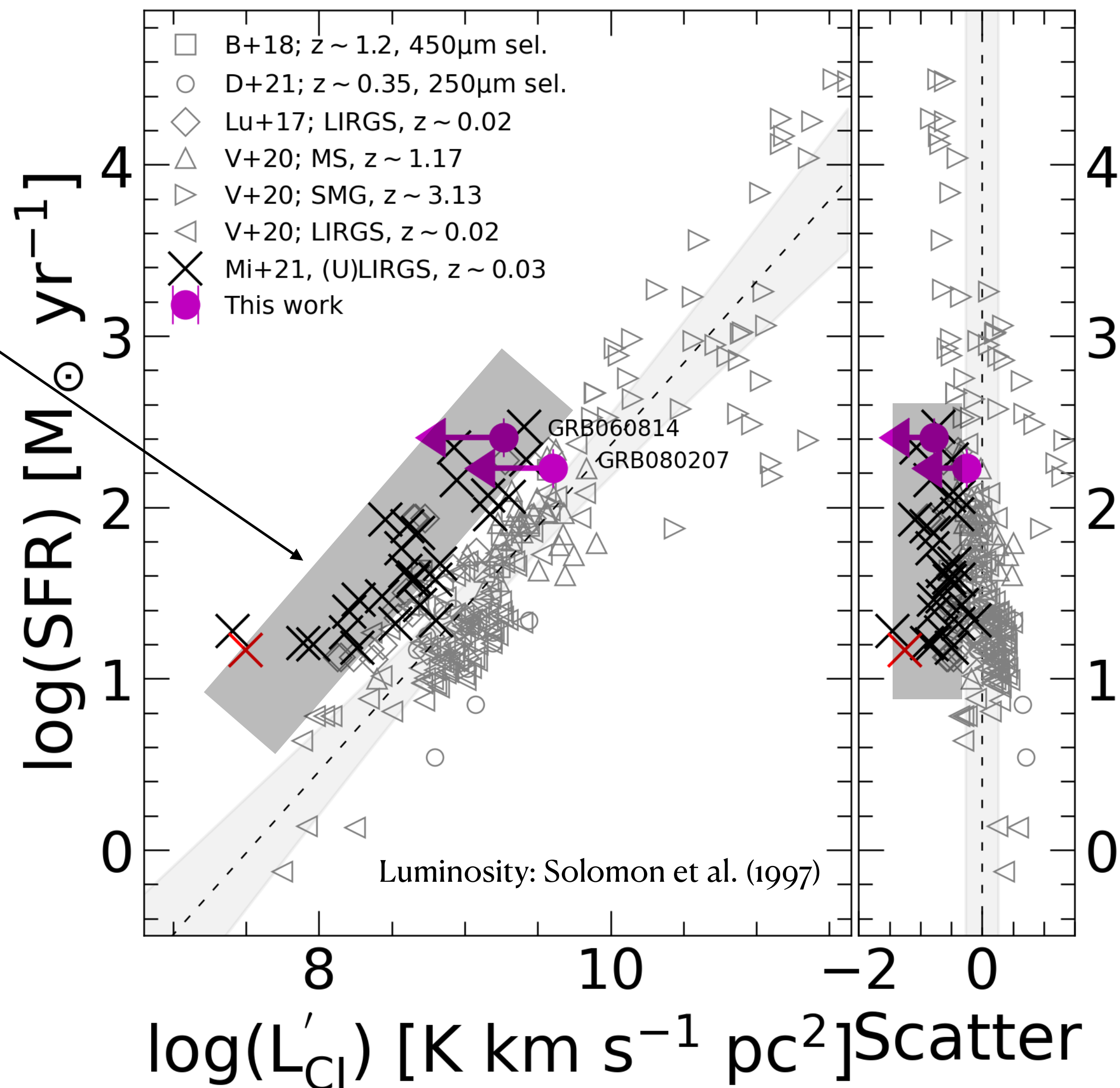
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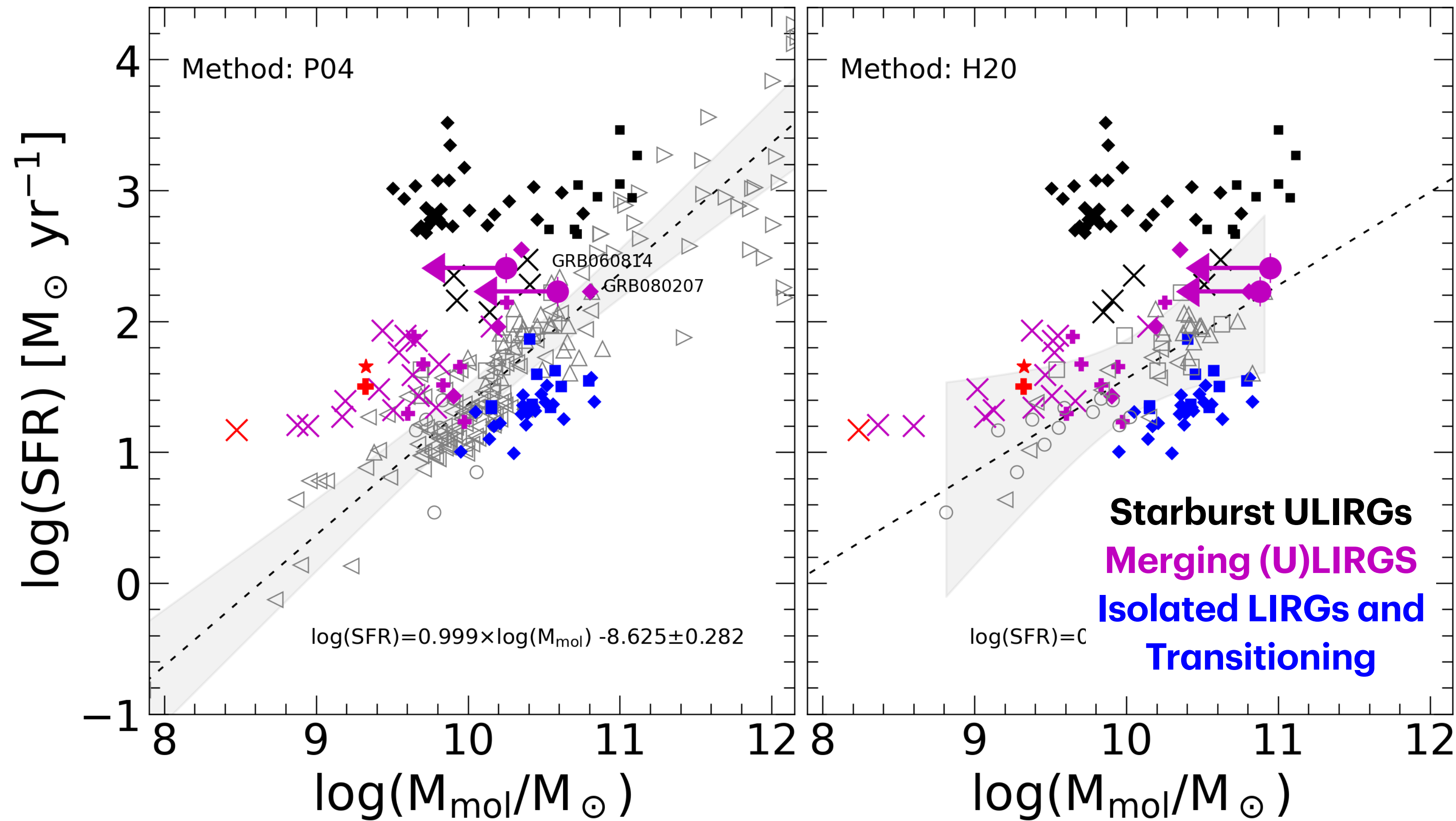
HST high-res images — *possible mergers*



Blanchard et al. (2016)



# Results



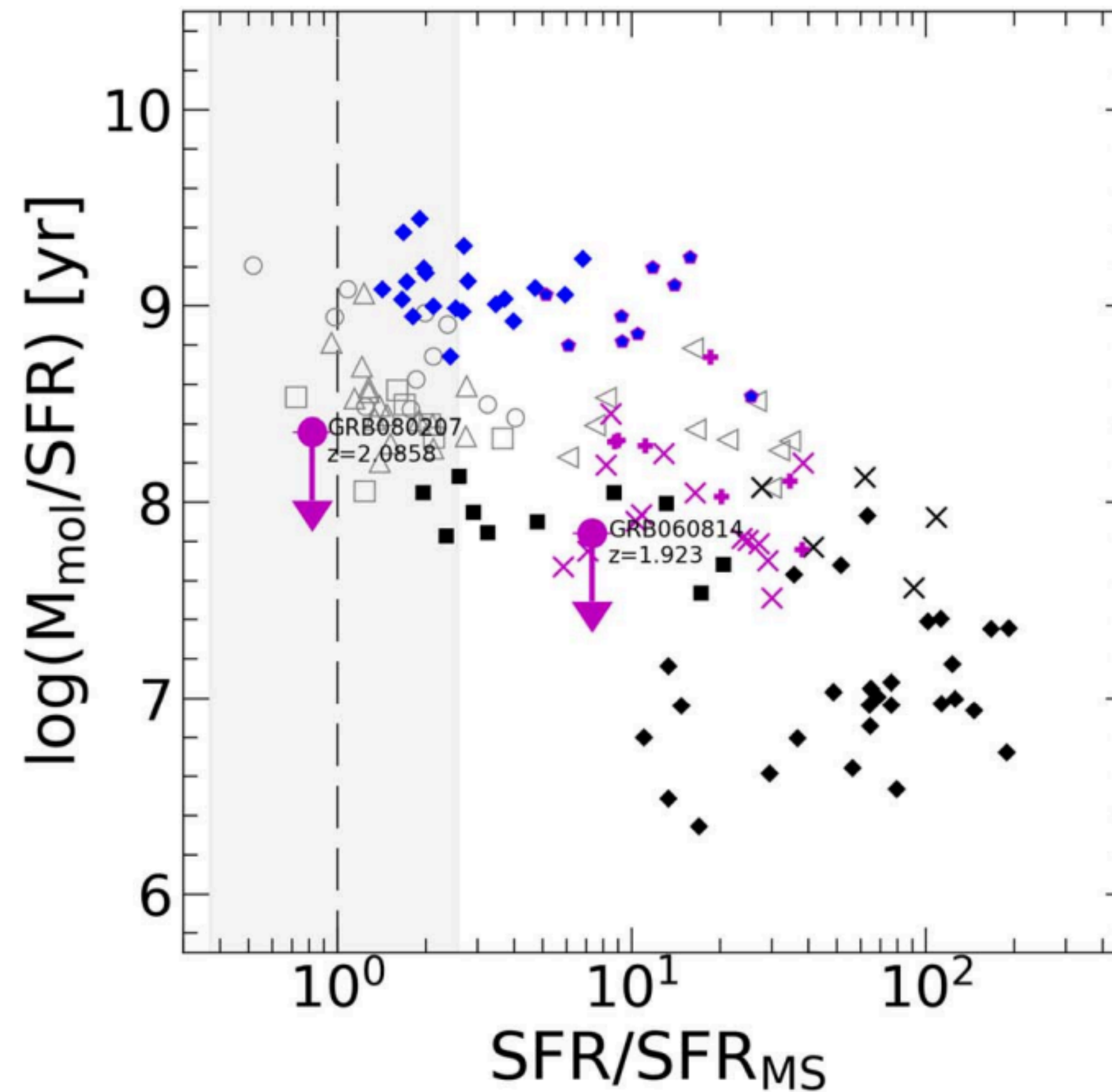
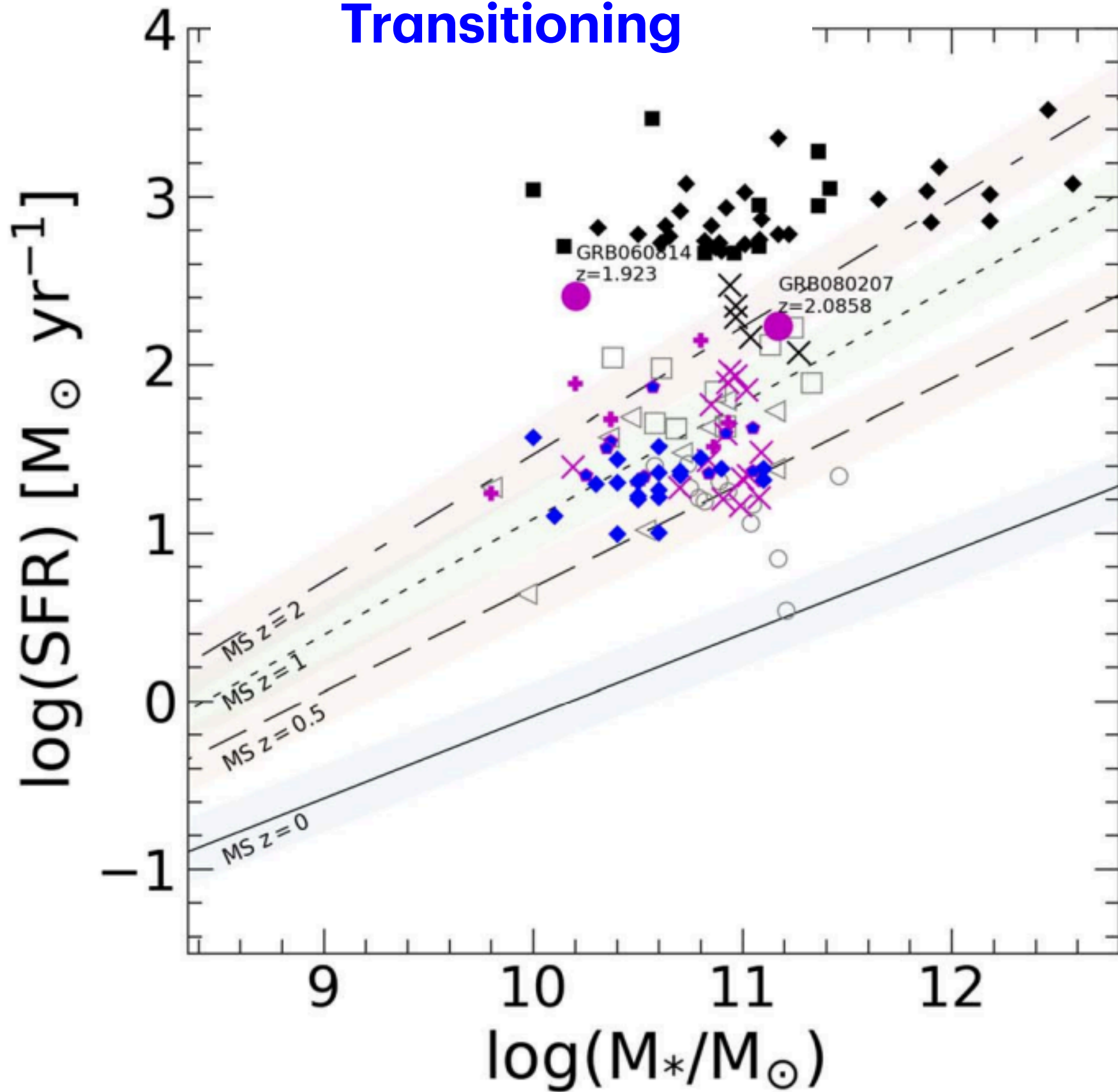
[CI] nondetections  $\rightarrow$  low [CI] mass (flux)

Independently of the method:  
For the same molecular mass  
we have a wide spread in SFR.

Molecular Masses: Papadopoulos et al. (2004); Heinz & Watson (2020)

# Results

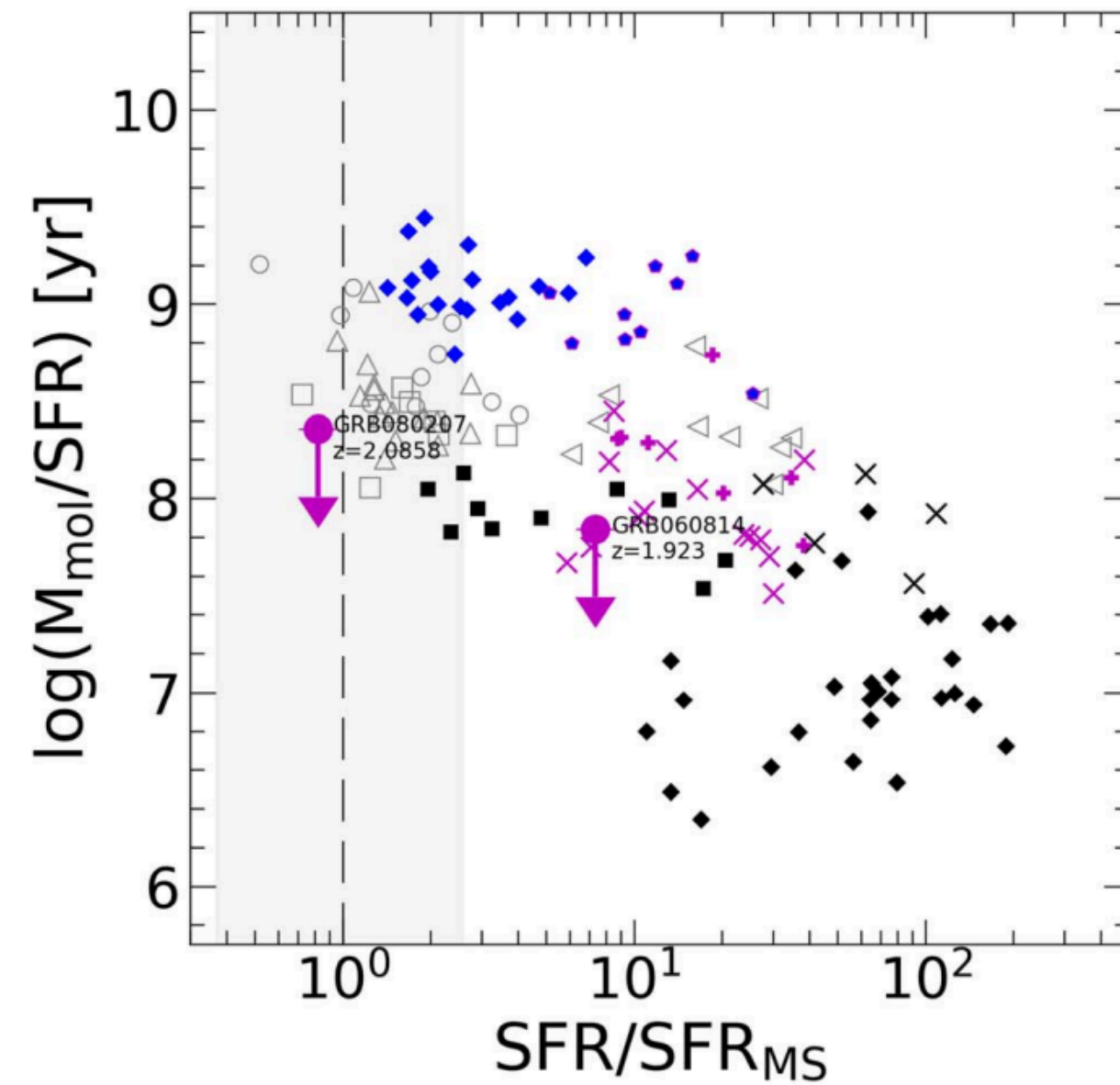
Starburst ULIRGs  
Merging (U)LIRGS  
Isolated LIRGs and  
Transitioning



Depletion time decrease with increasing distance from the MS. This correlates with a different stages found in galaxy evolution.

From isolated LIRGs, isolated transitioning, through transitioning merging ULIRGS, to starbursts.

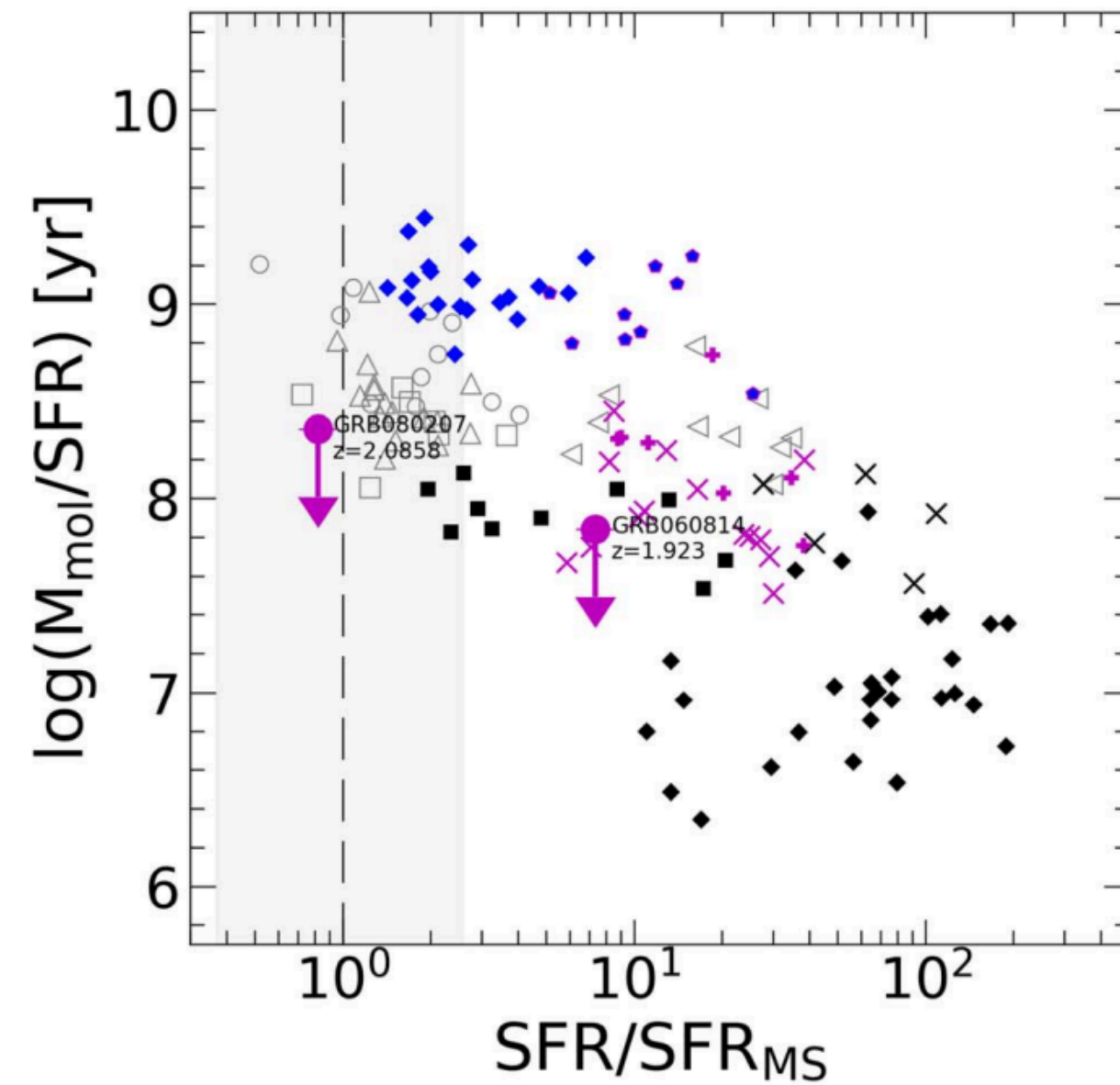
# Two nonexclusive scenarios



- 1) Our GRB hosts are at their transition toward starburst phase**  
(moderate distance from MS, short dep. times)
- 2) [CI]-dark objects** [Michiyama et al. \(2020\)](#)  
[CI]/CO < 0.1, from PDR models high H densities  
( $>10^5-10^6 \text{ cm}^{-3}$ )



# Two nonexclusive scenarios



**1) Our GRB hosts are at their transition toward starburst phase**

(moderate distance from MS, short dep. times)

**2) [CI]-dark objects** [Michiyama et al. \(2020\)](#)

[CI]/CO < 0.1, from PDR models high H densities (>10<sup>5</sup>-10<sup>6</sup> cm<sup>-3</sup>)

## and one rejected

**1) Our GRB hosts are in the post starburst phase**

Spectra of young stellar populations — no Balmer absorption lines

# Conclusions

**Using GRB as a tracer of recent star formation we found [CI]-dark galaxies that are in their transition from MS to starburst phase at  $z \sim 2$ .**

**We have no evidence for a difference between a GRB host and other SFGs.**

**Our GRB hosts are a high- $z$  analogues of local transitioning and [CI]-dark galaxies**

Read the paper here:

