



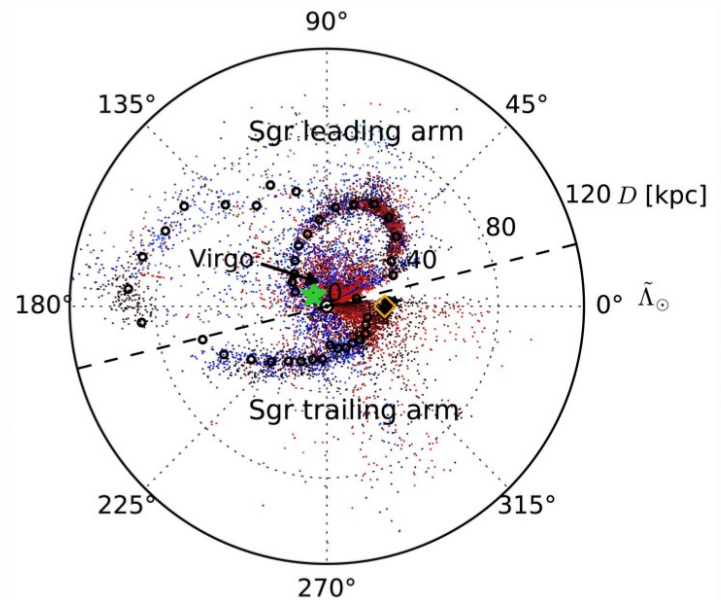
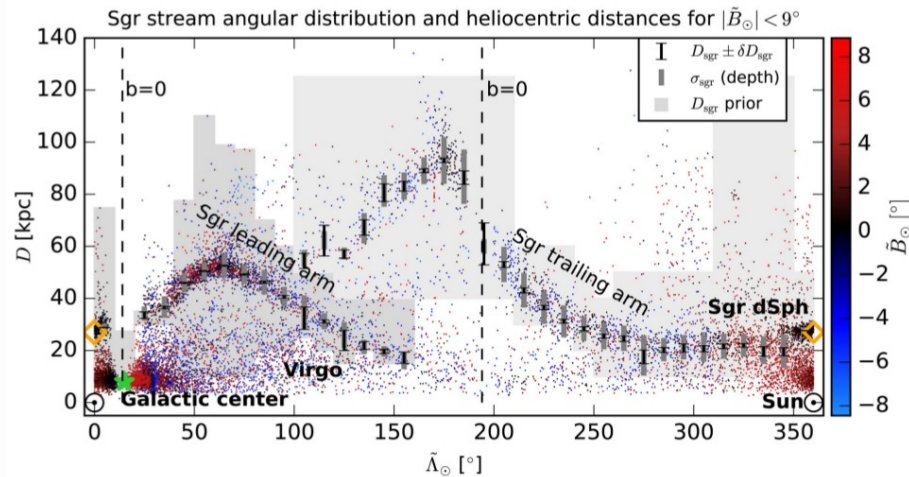
Circumstellar matter around RR Lyrae stars

Gergely Hajdu
Nicolaus Copernicus Astronomical Center,
Warsaw, Poland

RR Lyrae stars: excellent tracers of Population II

Easy to find:

- bright ($40 - 60 L_{\text{Sun}}$)
- characteristic light curve
- tracing old populations

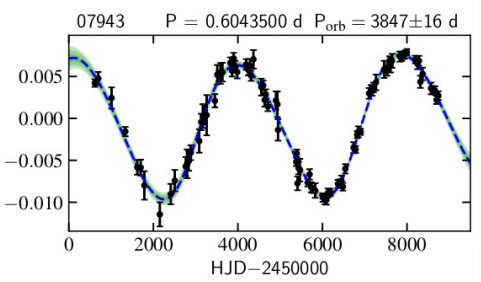
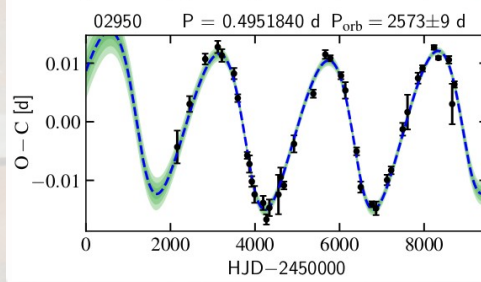
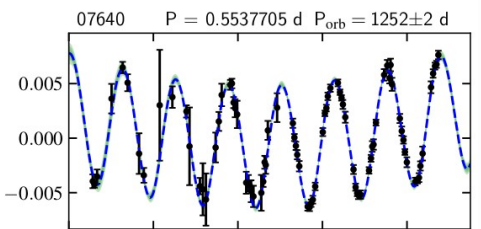
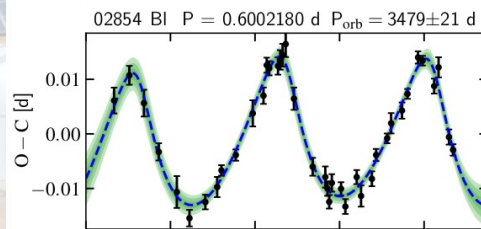
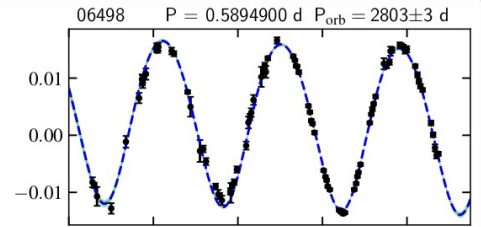
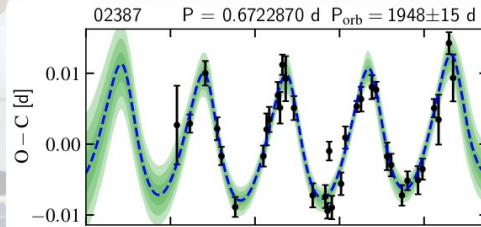


Hernitschek+2017

Search for binary RR Lyrae

Visually inspected:

- O-C of 27,480 OGLE RRab
- with LCs (folded and unfolded)
- 87 binary candidates (Hajdu+2021)



Search for binary RR Lyrae

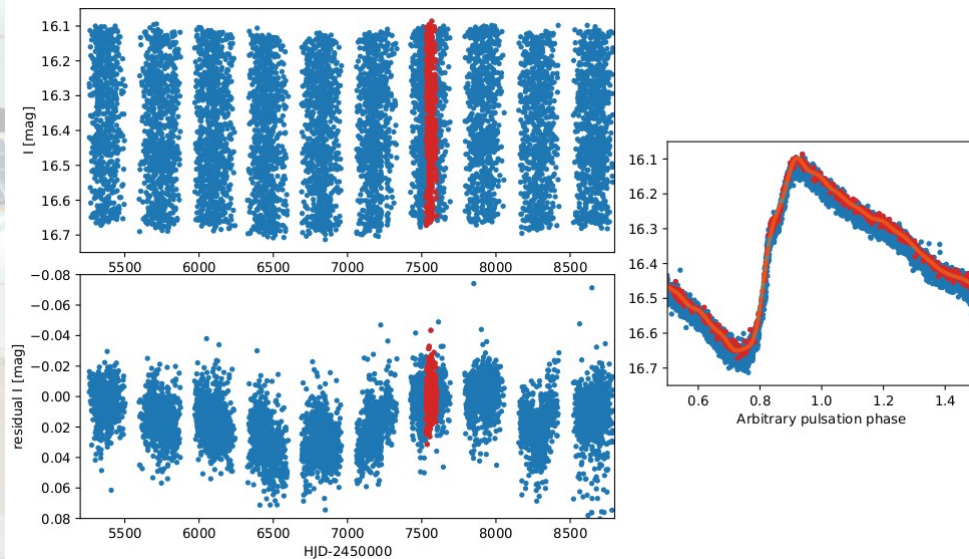
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Light-curve anomalies:

- some RR Lyrae change mean brightness with time

OGLE-BLG-RRLYR-33665



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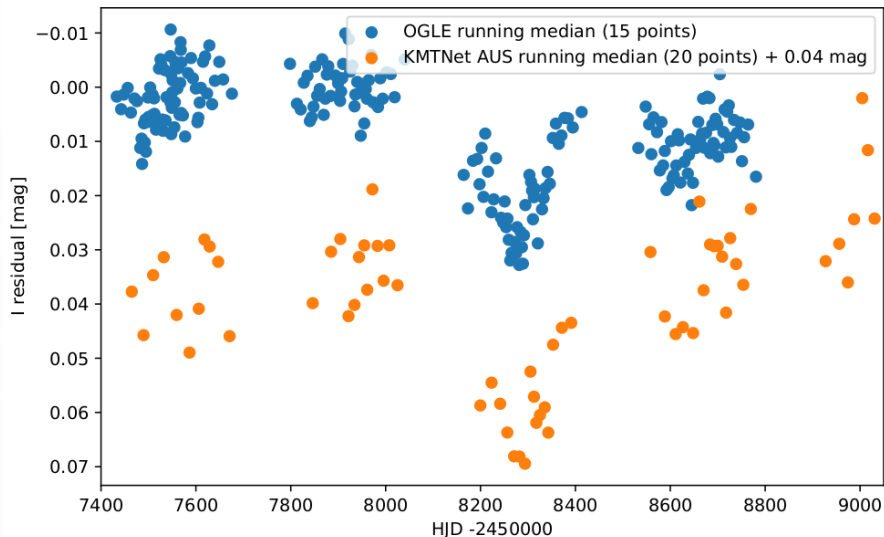
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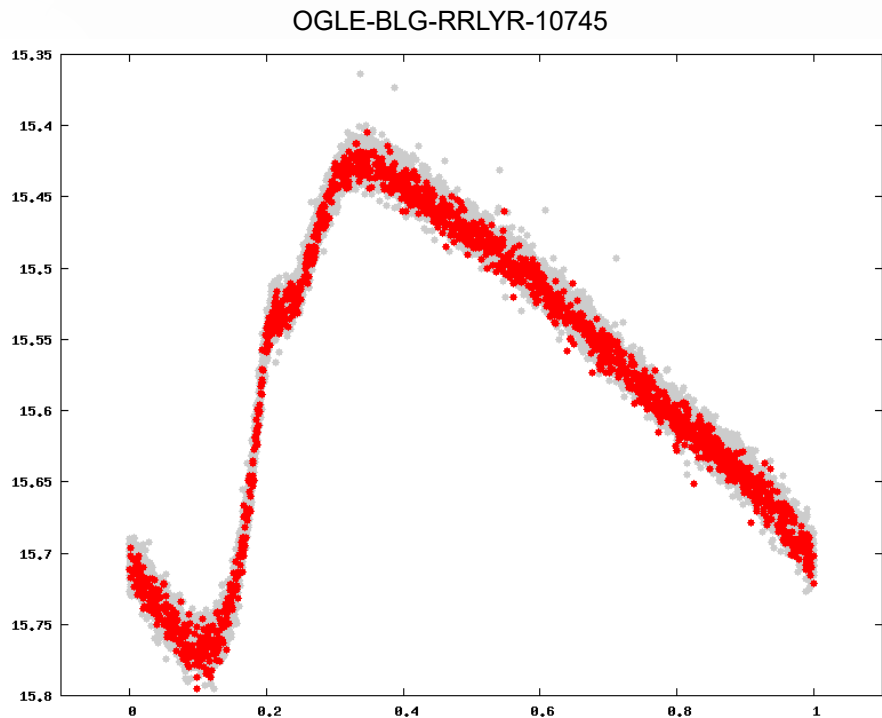
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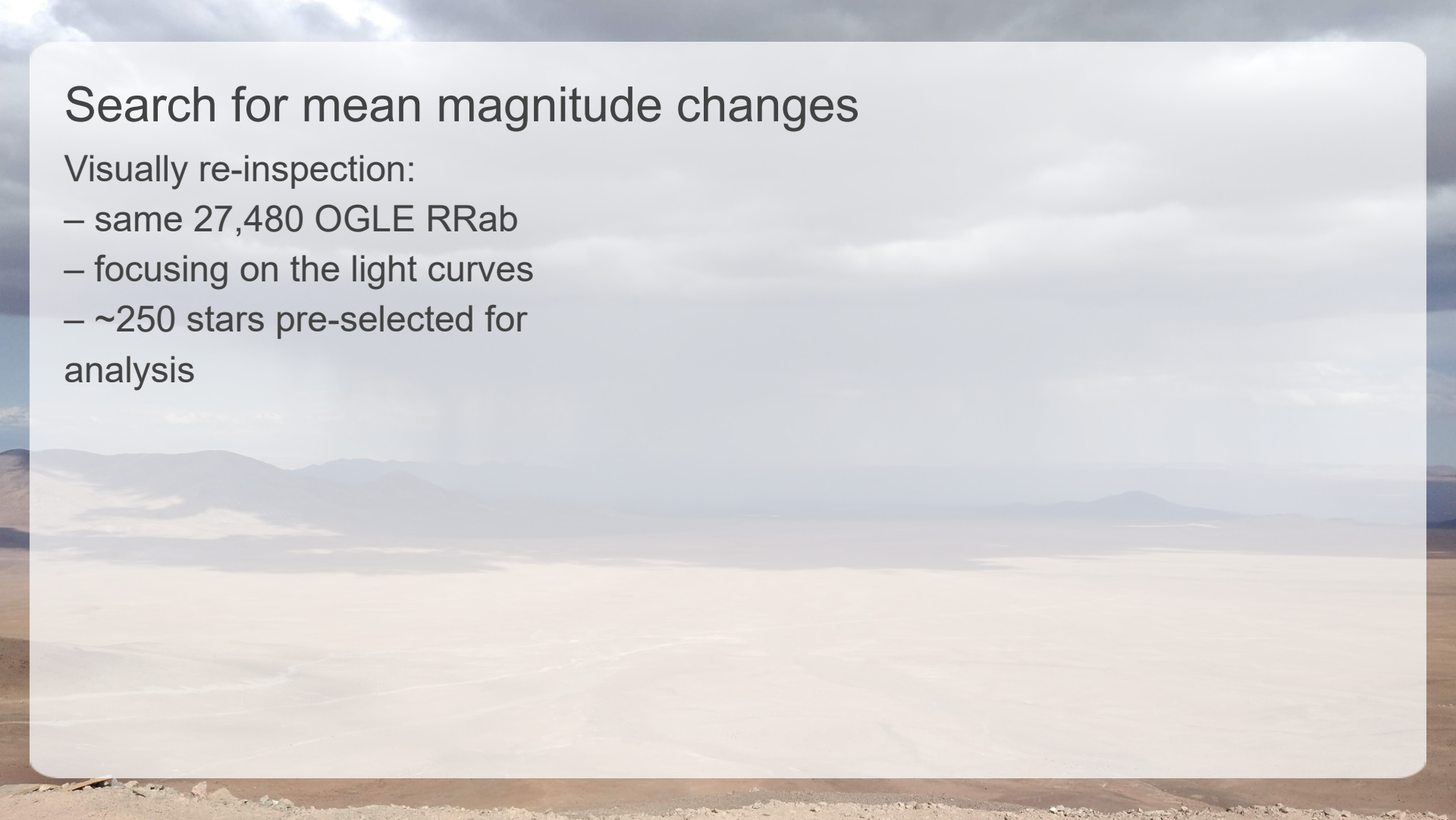
- some RR Lyrae change mean brightness with time
- some binary candidates too, and with the period of the binarity



Search for mean magnitude changes

Visually re-inspection:

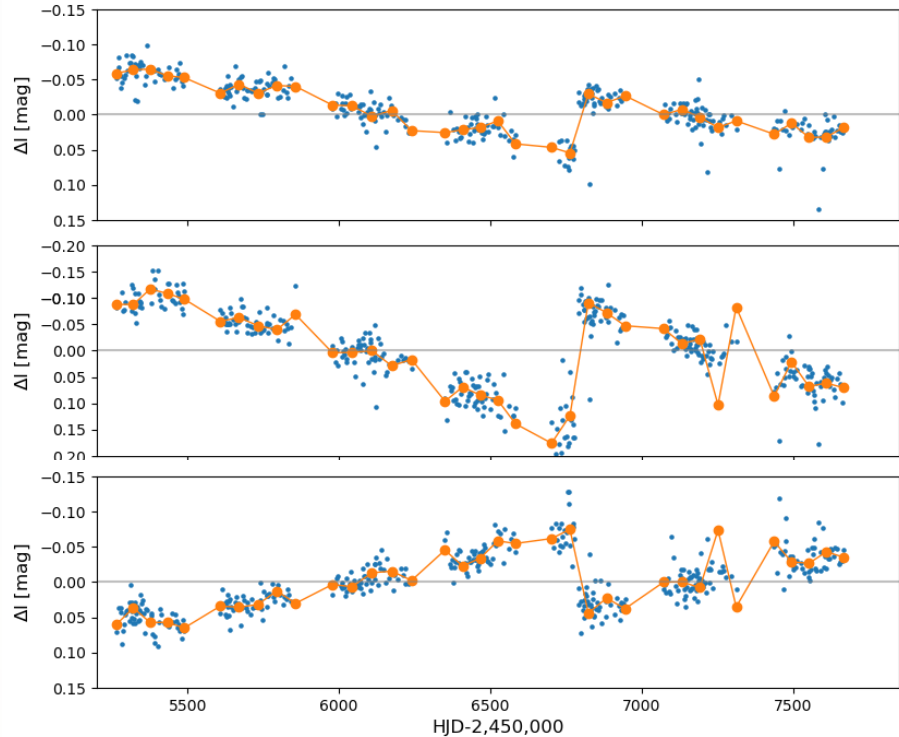
- same 27,480 OGLE RRab
- focusing on the light curves
- ~250 stars pre-selected for analysis



Search for mean magnitude changes

Visually re-inspection:

- same 27,480 OGLE RRab
 - focusing on the light curves
 - ~250 stars pre-selected for analysis
- analysis
- vetting out many false candidates
(blends, high proper motions,
many NGC 6441 RRL)

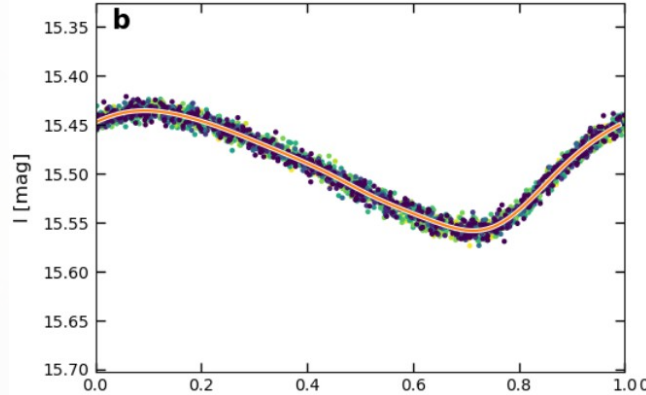
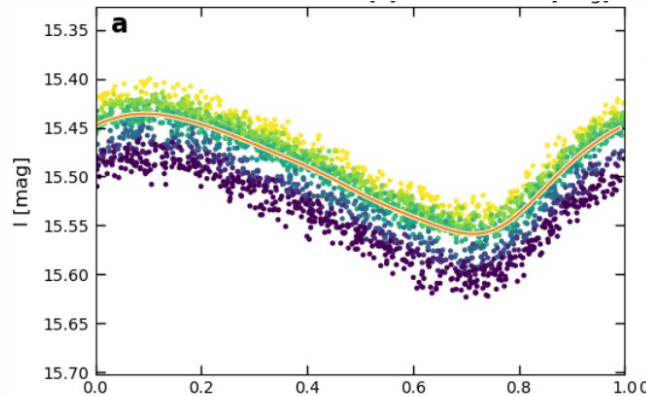


Modified Fourier fitting process

OGLE-BLG-RRLYR-12237

I-band fitting:

- Fourier series
- + variable mean magnitude



Modified Fourier fitting process

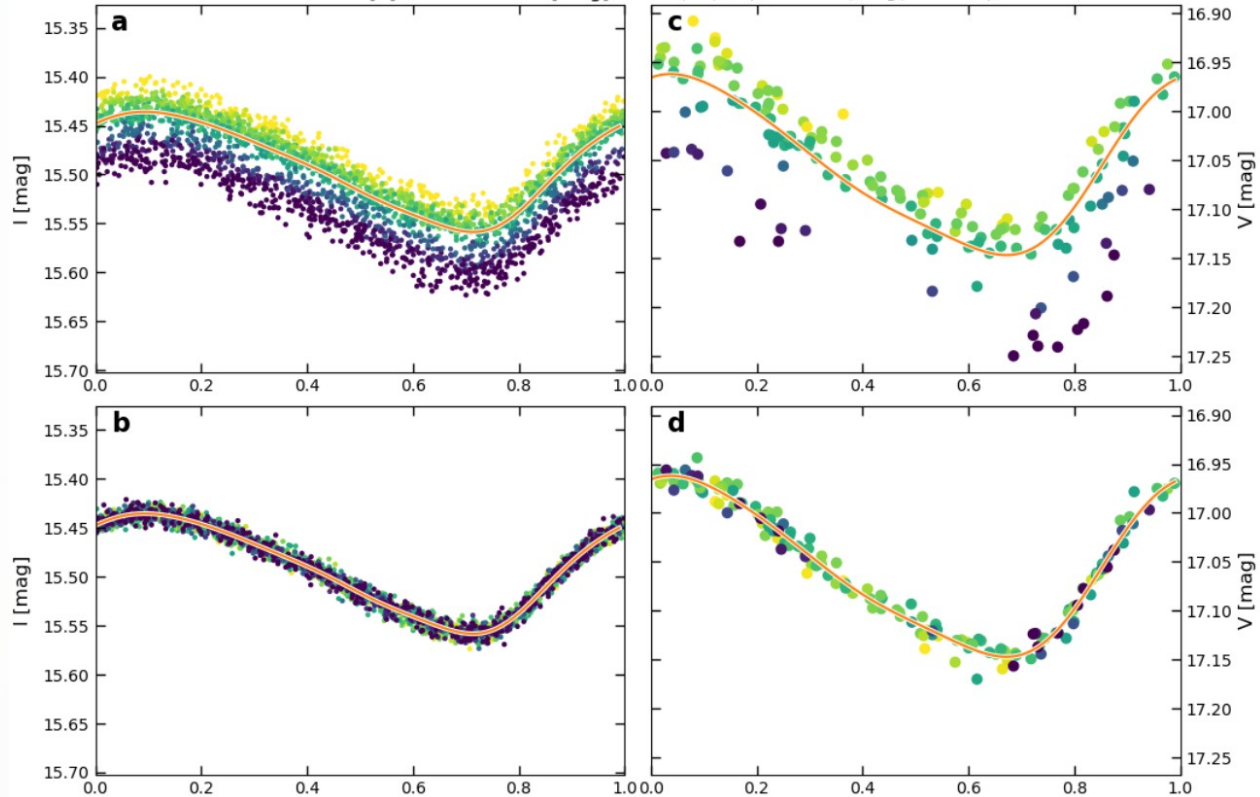
OGLE-BLG-RRLYR-12237

I-band fitting:

- Fourier series
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V-band fitting:

- Fourier series
- + I-band mean magnitude
- × extra constant



Modified Fourier fitting process

OGLE-BLG-RRLYR-12237

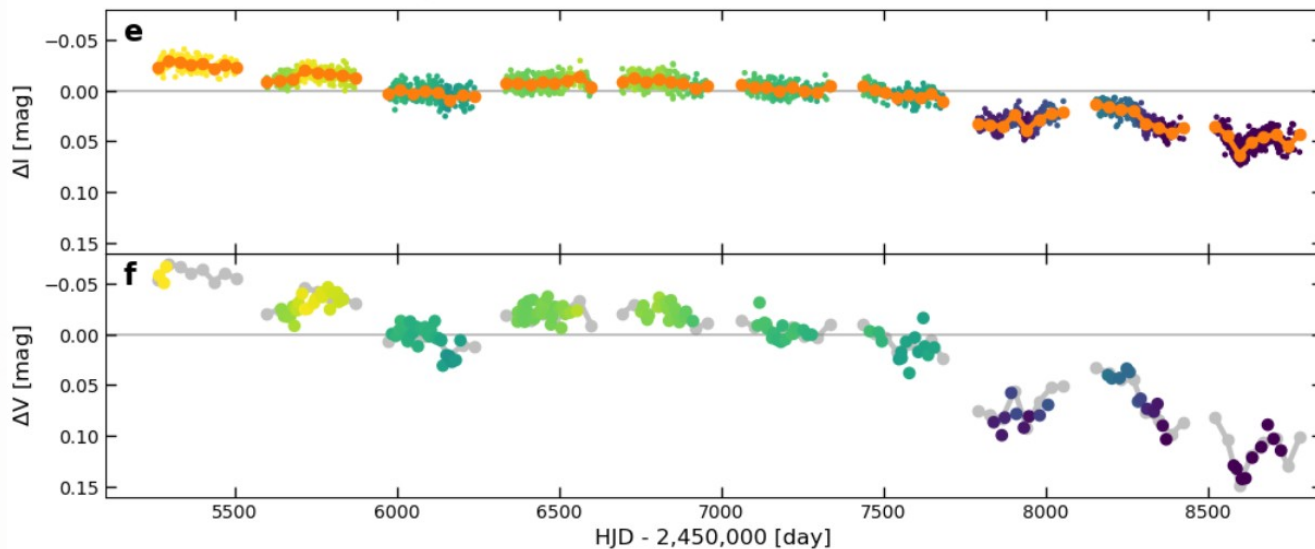
I-band fitting:

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V-band fitting:

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✗ extra constant
= $A(V) / A(I)$



Properties of the sample

Prevalence:

- 81 stars (80 RRab, 1 RRC)
- 5 / 81 binary candidates
- only a few stars with Blazhko effect



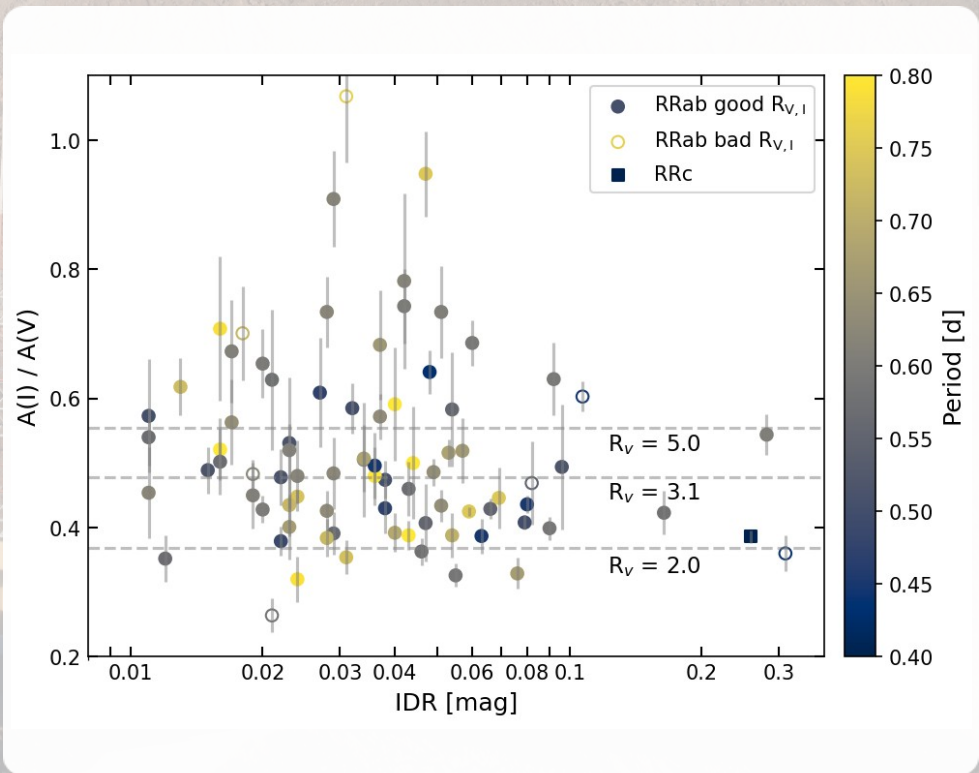
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A(I)/A(V) distribution:

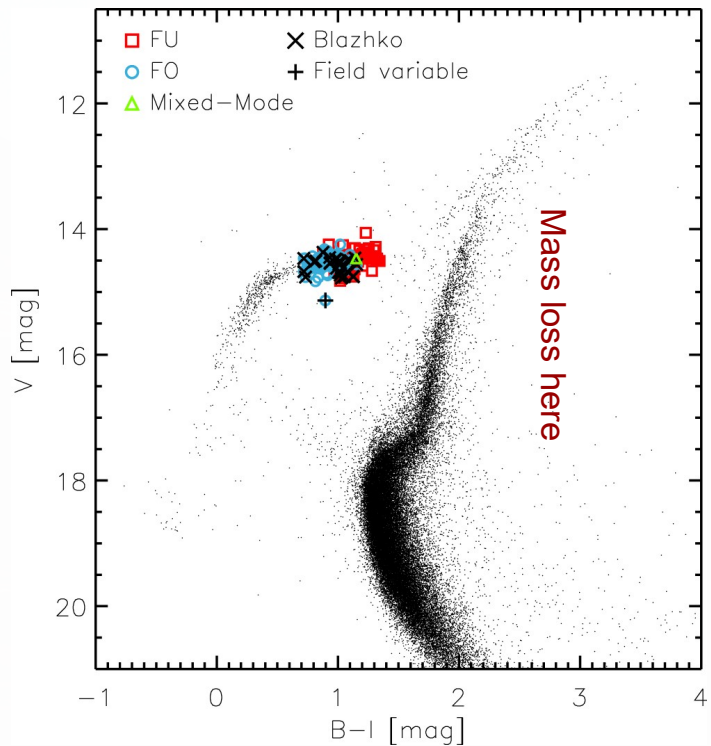
- very broad distribution from 0.35 to 0.8, outside normal interstellar extinction range
- supports a circumstellar (probably circumbinary) dusty origin



Connection to other kinds of stars

Things to consider:

- present in binary systems, $\sim P_{\text{orb}}$
- circumbinary dust? companion?
- source of dust? RRL itself? WD?



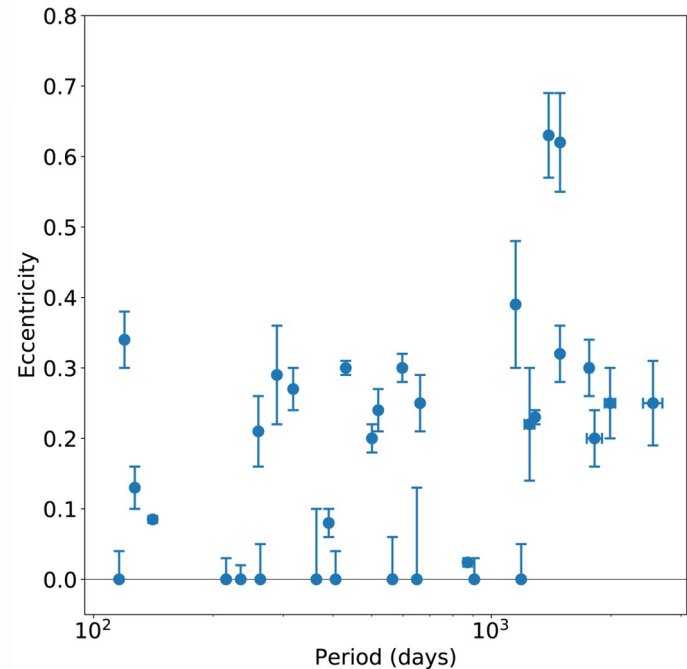
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Post-AGB binaries:

- $P_{\text{orb}} \sim 100\text{-}2500$ days
- circumbinary disks
- Lindblad resonance can pump the eccentricity (Glenn-Michael+2020)



Oomen+2018

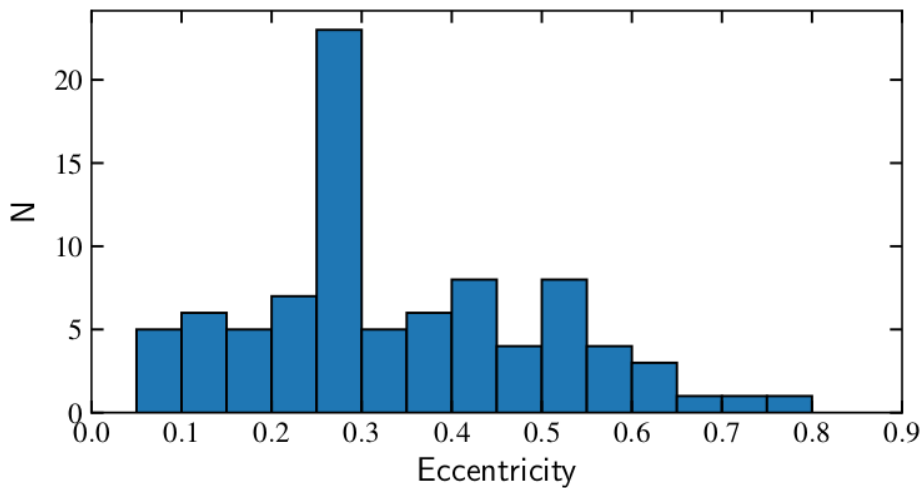
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- RR Lyrae: $e \sim 0.27$ excess



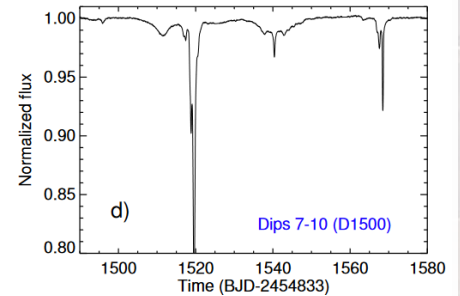
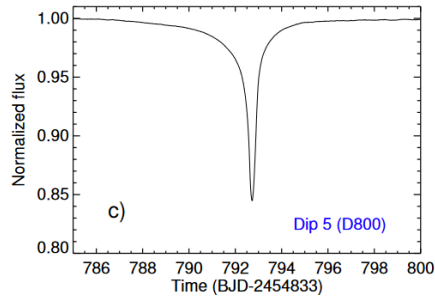
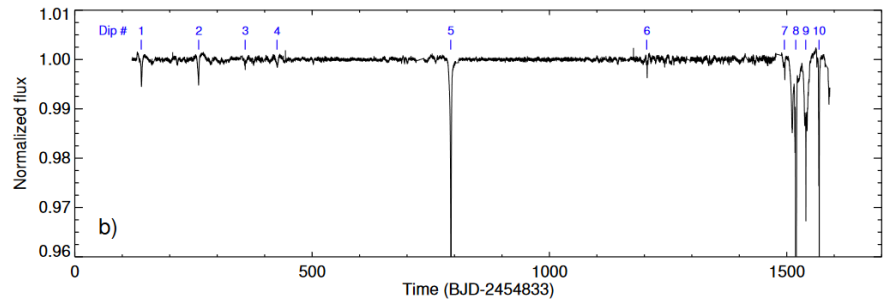
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KIC 8462852 (Boyajian's Star):

- mysterious dimmings in Kepler data



Boyajian+2015

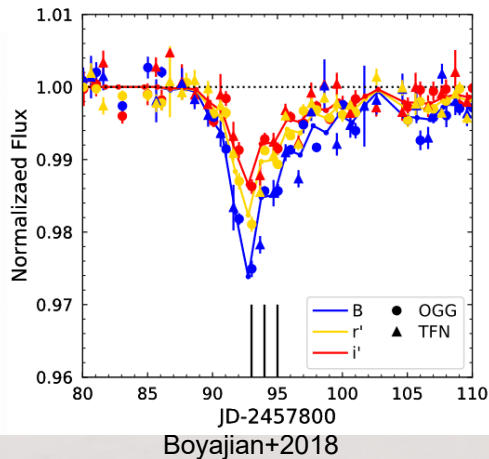
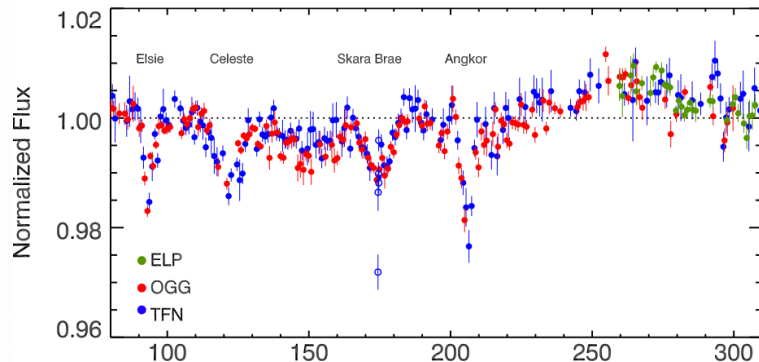
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- mysterious dimmings in Kepler data
- follow-up: new dips, chromatic behavior



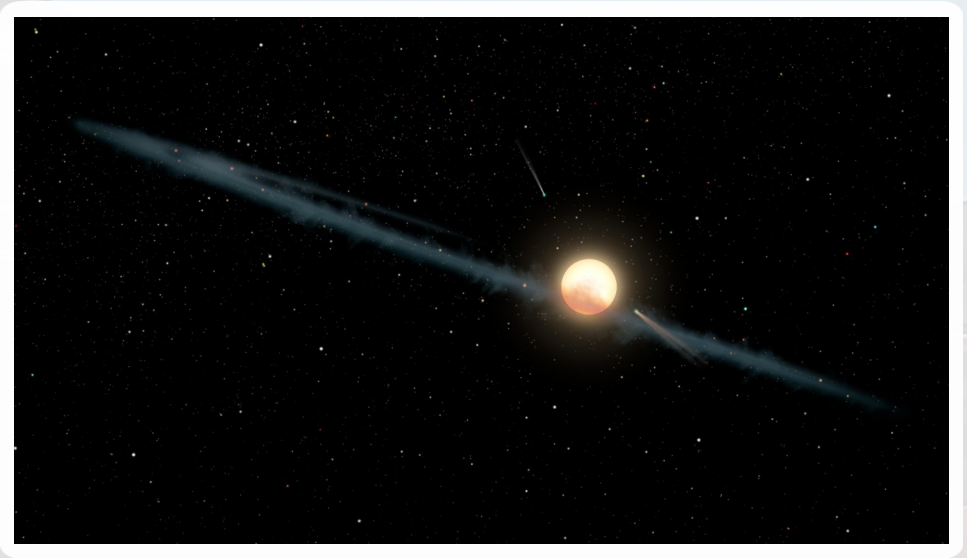
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KIC 8462852 (Boyajian's Star):

- mysterious dimmings in Kepler data
- follow-up: new dips, chromatic behavior
- uneven dust disk?



NASA/JPL-Caltech

Future prospects

Photometric studies needed:

- continued OGLE I-band, more V-band time series
- targeted multiband follow-up of the best stars + mid-IR
- other horizontal branch stars?



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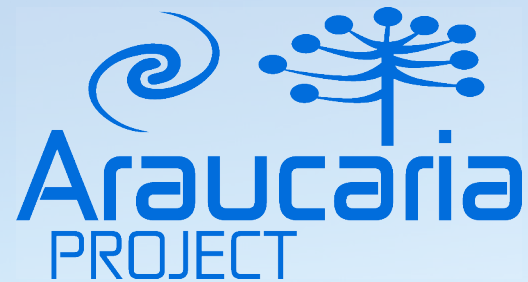
- continued OGLE I-band, more V-band time series
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Spectroscopy:

- chromatic Rossiter-McLaughlin effect might be present
- Ca H & K lines, other?

Polarimetry and simulation:

- scattered light might be detectable (time dependence)
- mass loss on RGB, binary evolution of RRL

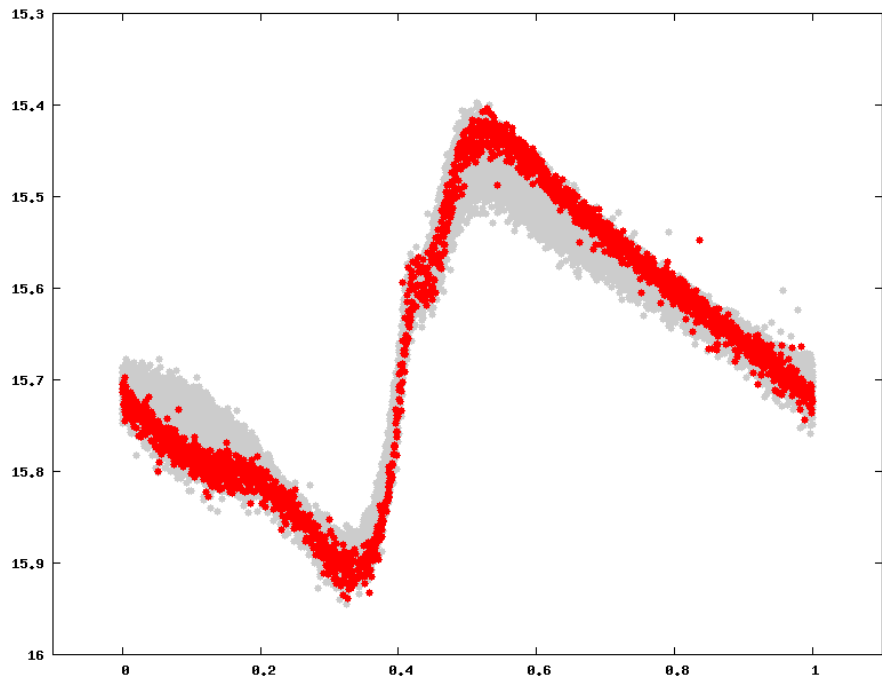


Thank you for your attention!

Some more curious RR Lyrae

OGLE-BLG-RRLYR-08752:

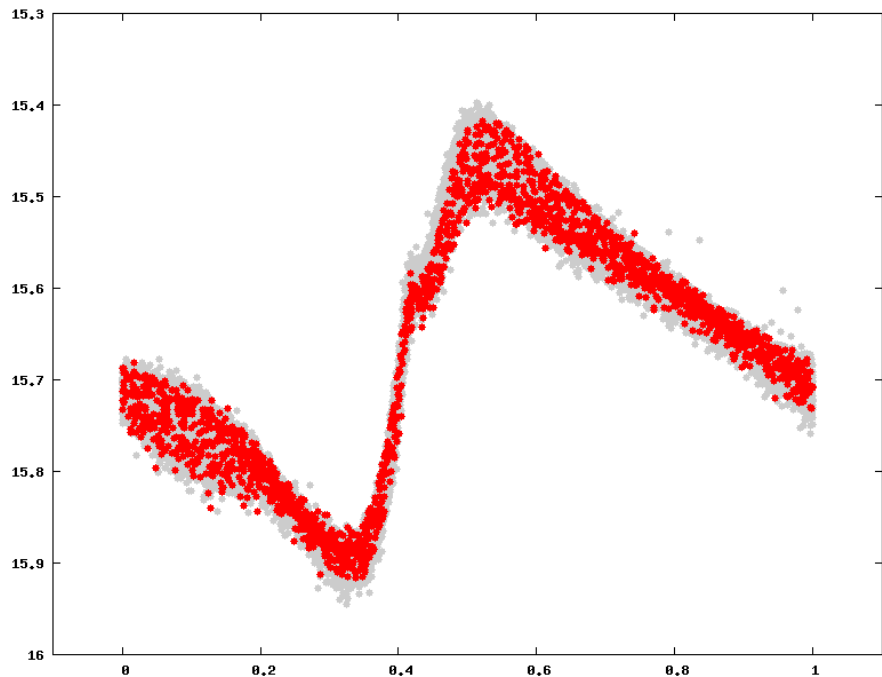
- binary candidate, 2 Blazhko modes
- + mean magnitude change (w/ P_{orb})



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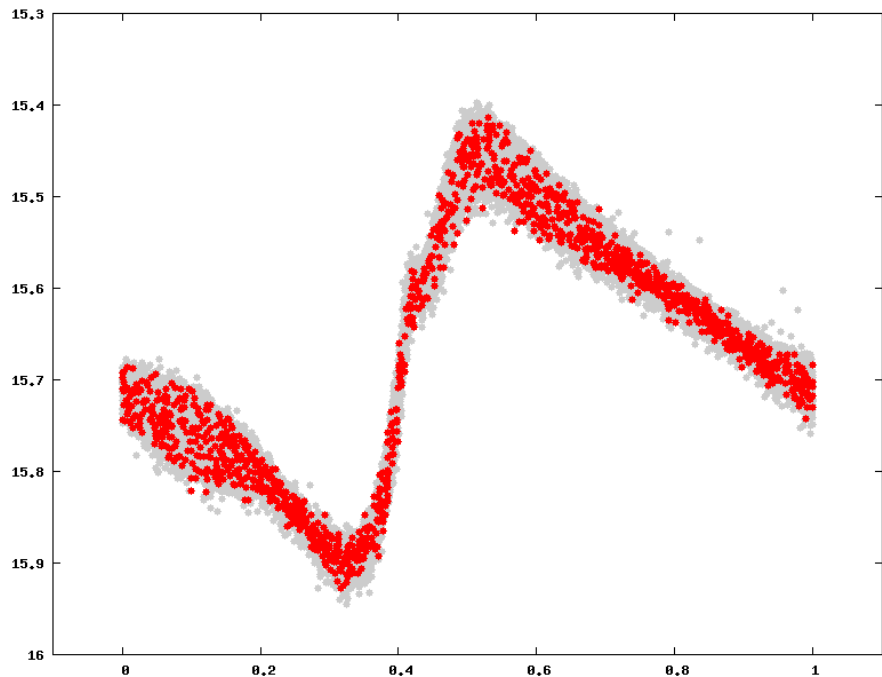
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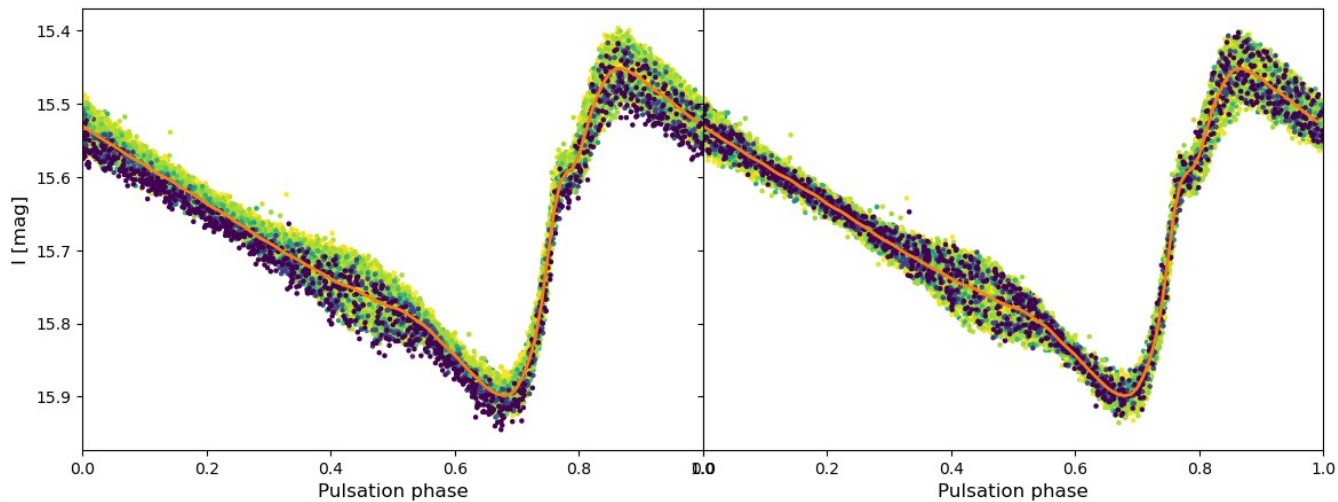
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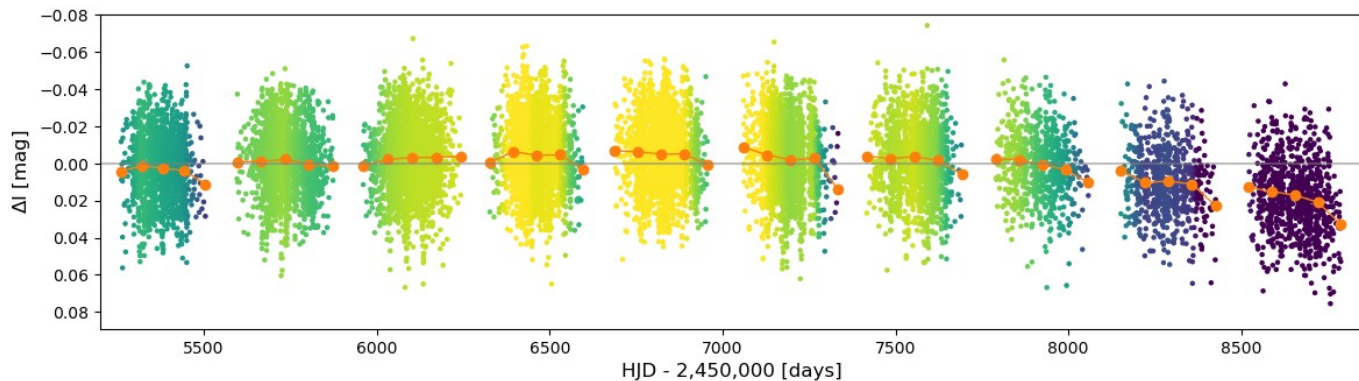
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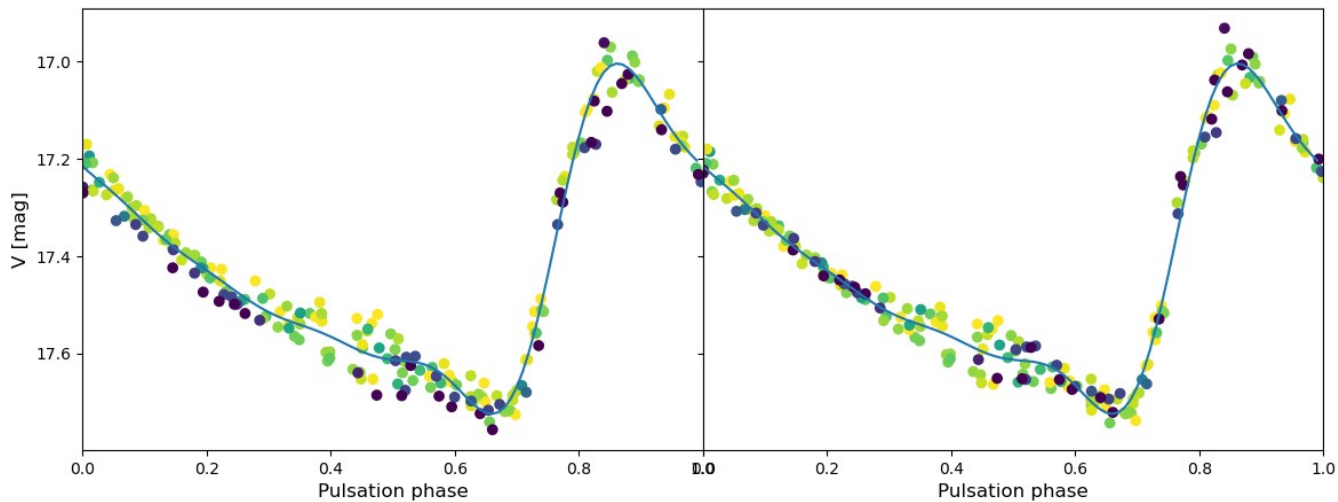
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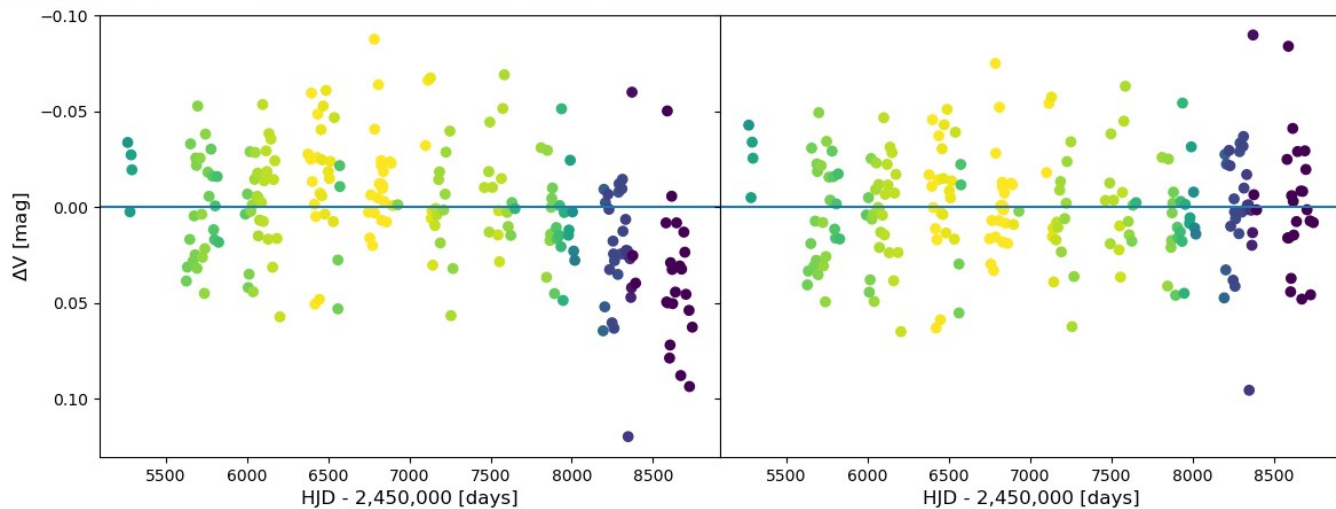
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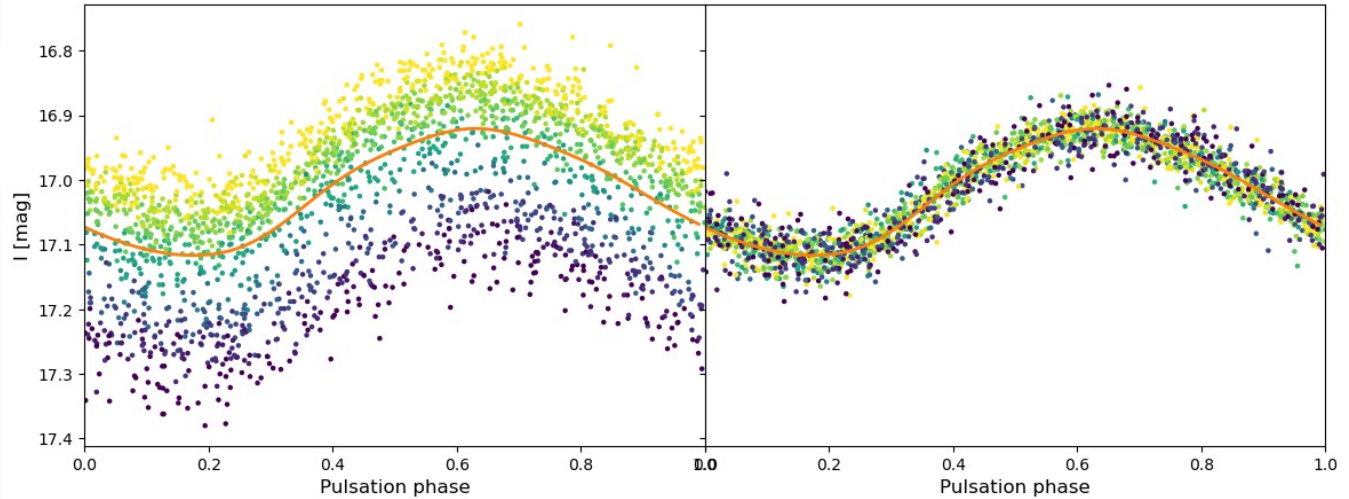
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OGLE-BLG-RRLYR-34373:

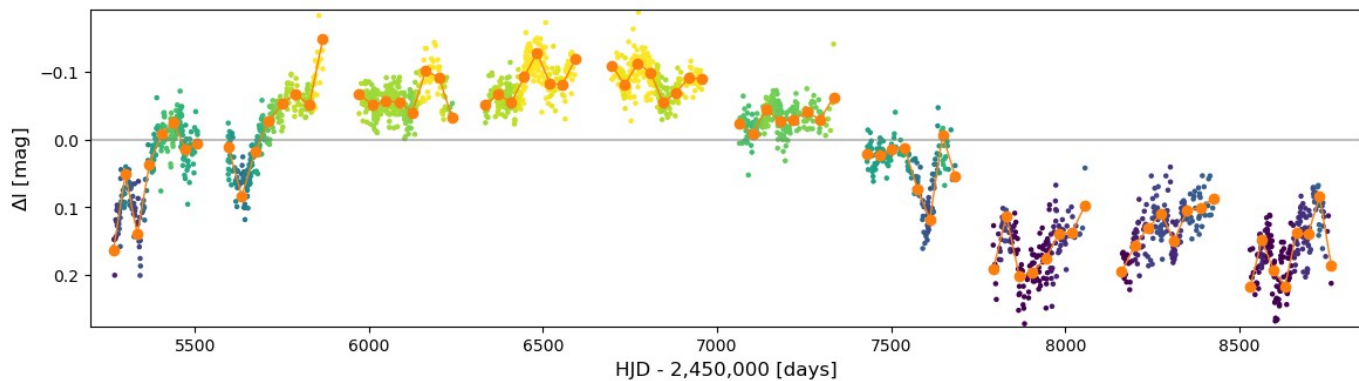
- RRc with variable mean magnitude noted in the OGLE catalog



Some more curious RR Lyrae

OGLE-BLG-RRLYR-34373:

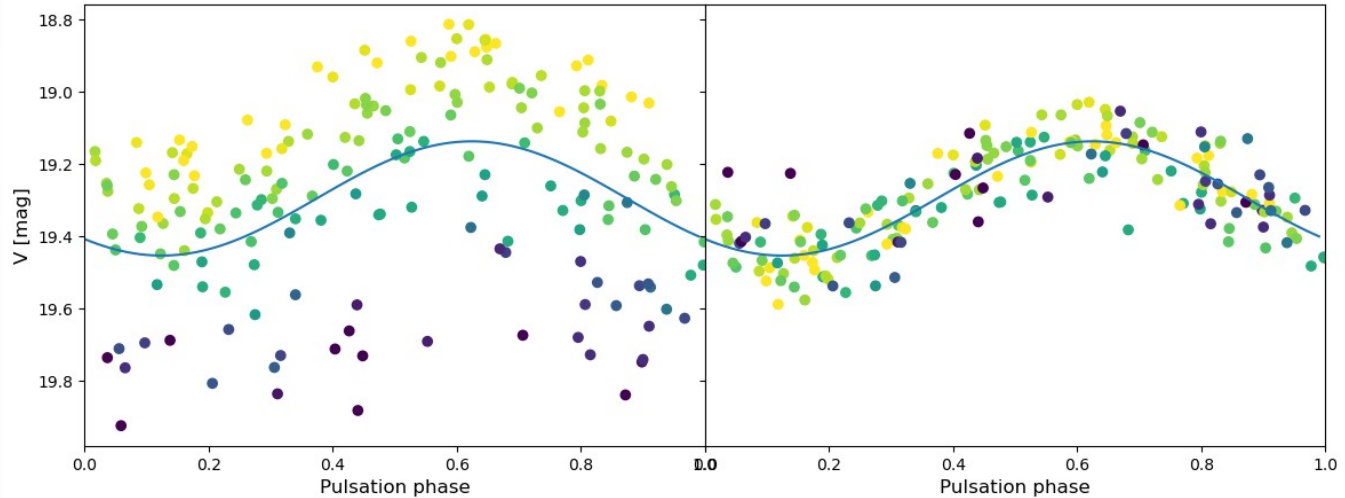
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